# Gibbs measures of nonlinear Schrödinger equations and many-body quantum mechanics

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## Classical mechanics and Gibbs measures

A Hamiltonian system consists of the following ingredients.

- Linear phase space  $\Gamma \ni \phi$ .
- Hamilton (or energy) function  $H \in C^{\infty}(\Gamma)$ .
- Poisson bracket  $\{\cdot, \cdot\}$  on  $C^{\infty}(\Gamma) \times C^{\infty}(\Gamma)$ .

(Properties: antisymmetric, bilinear, Leibnitz rule in both arguments, Jacobi identity.)

Classical dynamics is given by Hamiltonian flow  $\phi \mapsto S^t \phi$  on  $\Gamma$  defined by the ODE

$$\frac{\mathrm{d}}{\mathrm{d}t}f(S^t\phi) = \{H, f\}(S^t\phi)$$

for any  $f \in C^{\infty}(\Gamma)$ .

Standard example: classical system of n degrees of freedom.

• Phase space  $\Gamma = \mathbb{R}^{2n} \ni (p,q)$ .

Hamiltonian flow reads

$$\frac{\mathrm{d}}{\mathrm{d}t}p_i = -\frac{\partial H}{\partial q_i} = -\partial_i V(q), \qquad \frac{\mathrm{d}}{\mathrm{d}t}q_i = \frac{\partial H}{\partial p_i} = \frac{p_i}{m_i}$$

The Gibbs measure at temperature  $\beta$  is

$$\mathbb{P}(\mathrm{d}\phi) := \frac{1}{Z} \mathrm{e}^{-\beta H(\phi)} \,\mathrm{d}\phi \,, \qquad Z := \int \mathrm{e}^{-\beta H(\phi)} \,\mathrm{d}\phi \,.$$

 $\mathbb{P}$  is invariant under the flow  $S^t$ .

#### Nonlinear Schrödinger equations

Let  $\mathbb{T}^d = \mathbb{R}^d / \mathbb{Z}^d$  be the *d*-dimensional torus.

- Phase space  $\Gamma$  is some appropriate subspace of  $\{\phi : \mathbb{T}^d \to \mathbb{C}\}$ .
- Hamilton function

$$\begin{split} H(\phi) &= \int \mathrm{d}x\,\bar{\phi}(x)(\kappa-\Delta)\phi(x) + \frac{1}{2}\int \mathrm{d}x\,\mathrm{d}y\,w(x-y)|\phi(x)|^2|\phi(y)|^2\,,\\ \text{here }\kappa>0. \end{split}$$

Poisson bracket

W

$$\{\phi(x), \bar{\phi}(y)\} = \mathrm{i} \delta(x-y) \,, \qquad \{\phi(x), \phi(y)\} = \{\bar{\phi}(x), \bar{\phi}(y)\} = 0 \,.$$

Hamiltonian flow given by time-dependent nonlinear Schrödinger equation

$$\mathrm{i}\partial_t\phi(x) = (\kappa - \Delta)\phi(x) + \int \mathrm{d}y \, w(x - y) |\phi(y)|^2 \phi(x) \,.$$

Time-dependent nonlinear Schrödinger equation

$$i\partial_t \phi(x) = (\kappa - \Delta)\phi(x) + \int dy \, w(x - y) |\phi(y)|^2 \phi(x) \,. \tag{1}$$

Gibbs measure of nonlinear Schrödinger equation is formally

$$\mathbb{P}(\mathrm{d}\phi) = \frac{1}{Z} \mathrm{e}^{-H(\phi)} \mathrm{d}\phi \,.$$

Formally,  $\mathbb{P}$  is invariant under the flow generated by (1).

Rigorous results: Lebowitz–Rose–Speer, Bourgain, Bourgain–Bulut, Tzvetkov, Thomann–Tzvetkov, Nahmod–Oh–Rey-Bellet–Staffilani, Oh–Quastel, Deng–Tzvetkov–Visciglia, Cacciafesta–de Suzzoni, Genovese–Lucá–Valeri, ...

Important application:  $\mathbb P\text{-almost}$  sure well-posedness of (1) for rough initial data.

## Goal

#### Analyse $\mathbb{P}$ and $(S^t)$ through

- the moments of  $\mathbb{P}$  (which determine  $\mathbb{P}$ ),
- the time-dependent correlation functions

$$\int X^1(S^{t_1}\phi)\cdots X^m(S^{t_m}\phi)\,\mathrm{d}\mathbb{P}(\phi)\,,$$

for  $X^i \in C^{\infty}(\Gamma)$  and  $t_i \in \mathbb{R}$ .

Derivation as a (high-temperature) limit of a microscopic *n*-body quantum theory of bosons.

### Rigorous construction of Gibbs measure

Spectral decomposition

$$\kappa - \Delta = \sum_{k \in \mathbb{N}} \lambda_k u_k u_k^*, \qquad \lambda_k > 0, \qquad \|u_k\|_{L^2} = 1.$$

Let  $\omega = (\omega_k)_{k \in \mathbb{N}} \in \mathbb{C}^{\mathbb{N}}$  be i.i.d.  $\mathcal{N}_{\mathbb{C}}(0, 1)$  random variables with joint law  $\mu_0$ . Define the Gaussian free field

$$\phi^{\omega} \equiv \phi := \sum_{k \in \mathbb{N}} \frac{\omega_k}{\sqrt{\lambda_k}} u_k.$$

The sum converges in  $\|\phi\|_{\mathcal{H}^s} := \|(\kappa - \Delta)^{s/2}\phi\|_{L^2}$  in the sense of  $L^p(\mu_0)$  for all  $p \in (1, \infty)$ , provided that

$$\sum_{k\in\mathbb{N}}\lambda_k^{s-1}<\infty\,.$$

For example

e, 
$$\mathbb{E}^{\mu_0} \|\phi\|_{\mathcal{H}^s}^2 = \sum_{k \in \mathbb{N}} \mathbb{E}^{\mu_0} |\omega_k|^2 \frac{\lambda_k^s}{\lambda_k} = \sum_{k \in \mathbb{N}} \lambda_k^{s-1} \,.$$

Under  $\mu_0$ ,  $\phi = \sum_{k \in \mathbb{N}} \frac{\omega_k}{\sqrt{\lambda_k}} u_k$  is a Gaussian free field with covariance  $(\kappa - \Delta)^{-1}$ :

$$\mathbb{E}^{\mu_0} \langle f, \phi \rangle \langle \phi, g \rangle = \langle f, (\kappa - \Delta)^{-1} g \rangle.$$

We find that

$$\mu_0[\phi \in \mathcal{H}^0] = \begin{cases} 1 & \text{if } d = 1\\ 0 & \text{if } d > 1 \,. \end{cases}$$

Define the measure

$$\mu(d\omega) := \frac{1}{Z} e^{-W(\phi^{\omega})} \mu_0(d\omega), \qquad W(\phi) = \frac{1}{2} \int dx \, dy \, w(x-y) |\phi(x)|^2 |\phi(y)|^2$$

 $\boldsymbol{\mu}$  is well-defined for instance if

d = 1,  $w \in L^{\infty}$ , w positive definite,

since then  $0 \leq W(\phi) < \infty \mu_0$ -a.s. (Then  $\mathbb{P}$  is defined as  $\phi_*\mu$ .)

### Quantum many-body theory

Define the one-particle space  $\mathfrak{H}:=L^2(\mathbb{T}^d;\mathbb{C})$  and the n-particle space

$$\mathfrak{H}^{(n)} := \mathfrak{H}^{\otimes_{\mathrm{sym}} n} = L^2_{\mathrm{sym}} \big( (\mathbb{T}^d)^n \big) \,.$$

Hamilton operator

$$H^{(n)} := H_0^{(n)} + \lambda \sum_{1 \le i < j \le n} w(x_i - x_j), \qquad H_0^{(n)} := \sum_{i=1}^n (\kappa - \Delta_{x_i})$$

Canonical thermal state at temperature  $\tau > 0$  is  $P_{\tau}^{(n)} := e^{-H^{(n)}/\tau}$ . Expectation of an observable  $A \in \mathfrak{B}(\mathfrak{H}^{(n)})$  is

$$\rho_{\tau}^{(n)}(A) := \frac{\operatorname{Tr}(AP_{\tau}^{(n)})}{\operatorname{Tr}(P_{\tau}^{(n)})}$$

What happens as  $n \to \infty$ ?

In order to obtain a nontrivial limit, we set  $\lambda = 1/n$ .

**Theorem** [Lewin-Nam-Serfaty-Solovej, 2012; Lewin-Nam-Rougerie, 2013]. For  $\lambda = 1/n$  and  $\tau$  fixed, the state  $\rho_{\tau}^{(n)}(\cdot)$  converges to the atomic measure  $\delta_{\Phi}$  in the sense of *p*-particle correlation functions (see later), where  $\Phi$  is the minimizer of the energy function *H*.

Complete Bose-Einstein condensation for fixed  $\tau$ .

In order to obtain the Gibbs measure  $\mu$ , we need to let

- $\tau$  grow with n (high-temperature limit),
- *n* fluctuate.  $(n/\tau \text{ will correspond to } \|\phi\|_2^2.)$

#### High-temperature limit and Fock space

Define the Fock space  $\mathcal{F}:=igoplus_{n\in\mathbb{N}}\mathfrak{H}^{(n)}$  and the grand canonical thermal state

$$P_{\tau} := \bigoplus_{n \in \mathbb{N}} P_{\tau}^{(n)} = e^{-H_{\tau}}, \qquad H_{\tau} := \frac{1}{\tau} \bigoplus_{n \in \mathbb{N}} H^{(n)}$$

Rescaled particle number operator  $\mathcal{N}_{\tau} := \frac{1}{\tau} \bigoplus_{n \in \mathbb{N}} nI$ . Expectation of an observable  $A \in \mathfrak{B}(\mathcal{F})$  is

$$o_{\tau}(A) := rac{\operatorname{Tr}(AP_{\tau})}{\operatorname{Tr}(P_{\tau})}.$$

Explicit computation for d = 1 and  $\lambda = 0$ :

$$\lim_{\tau \to \infty} \rho_{\tau}(\mathcal{N}_{\tau}^k) = \mathbb{E}^{\mu} \|\phi\|_{L^2}^{2k}, \qquad k = 1, 2, \dots.$$

Number of particles is of order  $\tau$ . Thus, set  $\lambda := \tau^{-1}$  to obtain nontrivial interacting limit.

#### Second quantization

Let  $b,b^*$  be the bosonic annihilation and creation operators on  ${\mathcal F}$  and set  $\phi:=\tau^{-1/2}b.$  Hence,

$$[\phi_{\tau}(x), \phi_{\tau}^{*}(y)] = \frac{1}{\tau} \delta(x - y), \qquad [\phi_{\tau}(x), \phi_{\tau}(y)] = [\phi_{\tau}^{*}(x), \phi_{\tau}^{*}(y)] = 0.$$

Thus, we can write  $H_{ au} = H_{ au,0} + W$ , where

$$H_{\tau,0} = \int \mathrm{d}x \,\phi_{\tau}^*(x)(\kappa - \Delta)\phi_{\tau}(x) \,,$$
$$W_{\tau} = \frac{1}{2} \int \mathrm{d}x \,\mathrm{d}y \,\phi_{\tau}^*(x)\phi_{\tau}(x) \,w(x - y) \,\phi_{\tau}^*(y)\phi_{\tau}(y) \,,$$

as well as  $P_{\tau} = e^{-H_{\tau}}$ ,  $\rho_{\tau}(A) = \frac{\operatorname{Tr}(AP_{\tau})}{\operatorname{Tr}(P_{\tau})}$ .

## High-temperature limit for d = 1

Define the *p*-particle reduced density matrix

$$\gamma_{\tau,p}(x_1,\ldots,x_p;y_1,\ldots,y_p) := \rho_\tau \left( \phi_\tau^*(y_1)\cdots \phi_\tau^*(y_p)\phi_\tau(x_1)\cdots \phi_\tau(x_p) \right).$$

Analogously, we define the classical p-particle correlation function

$$\gamma_p(x_1,\ldots,x_p;y_1,\ldots,y_p) := \mathbb{E}^{\mu}\big(\bar{\phi}(y_1)\cdots\bar{\phi}(y_p)\phi(x_1)\cdots\phi(x_p)\big)$$

The family  $(\gamma_p)_{p \in \mathbb{N}}$  completely determines all moments of the field  $\phi$ .

**Theorem** [Lewin-Nam-Rougerie, 2015]. For d = 1 and w positive definite, for any  $p \in \mathbb{N}$  we have  $\gamma_{\tau,p} \to \gamma_p$  in trace class as  $\tau \to \infty$ .

#### Time-dependent correlations

Introduce Hamiltonian time evolution:

(CI) for a random variable  $X \equiv X(\phi)$  set  $\Psi^t X(\phi) := X(S^t \phi)$ ; (Qu) for an operator **X** on  $\mathcal{F}$  set  $\Psi^t_{\tau} \mathbf{X} := e^{it\tau H_{\tau}} \mathbf{X} e^{-it\tau H_{\tau}}$ .

For a p-particle operator  $\xi$  on  $\mathfrak{H}^{(p)}$  introduce the observables

(CI) 
$$\Theta(\xi) := \int dx_1 \cdots dx_p dy_1 \cdots dy_p \xi(x_1, \dots, x_p; y_1, \dots, y_p)$$
  
  $\times \bar{\phi}(x_1) \cdots \bar{\phi}(x_p) \phi(y_1) \cdots \phi(y_p);$ 

$$(\mathsf{Qu}) \quad \Theta_{\tau}(\xi) := \int \mathrm{d}x_1 \cdots \mathrm{d}x_p \, \mathrm{d}y_1 \cdots \mathrm{d}y_p \, \xi(x_1, \dots, x_p; y_1, \dots, y_p) \\ \times \phi_{\tau}^*(x_1) \cdots \phi_{\tau}^*(x_p) \phi_{\tau}(y_1) \cdots \phi_{\tau}(y_p) \, .$$

**Theorem** [Fröhlich-K-Schlein-Sohinger, 2017]. Let d = 1 and  $w \in L^{\infty}$  be pointwise positive. Given  $m \in \mathbb{N}$ ,  $p_1, \ldots, p_m \in \mathbb{N}$ ,  $\xi^1 \in \mathcal{L}(\mathfrak{H}^{(p_1)}), \ldots, \xi^m \in \mathcal{L}(\mathfrak{H}^{(p_m)})$  and  $t_1, \ldots, t_m \in \mathbb{R}$ , we have

 $\lim_{\tau \to \infty} \rho_{\tau} \left( \Psi_{\tau}^{t_1} \Theta_{\tau}(\xi^1) \cdots \Psi_{\tau}^{t_m} \Theta_{\tau}(\xi^m) \right) = \mathbb{E}^{\mu} \left( \Psi^{t_1} \Theta(\xi^1) \cdots \Psi^{t_m} \Theta(\xi^m) \right).$ 

Remarks:

- Also works on  $\mathbb{R}$  instead of  $\mathbb{T}$ , with sufficiently confining potential v in free Hamiltonian  $\kappa \Delta + v(x)$ .
- We also prove that there exists a null sequence  $\varepsilon = \varepsilon_{\tau}$  such that, with a quantum two-body potential  $\frac{1}{\varepsilon}w(\frac{x}{\varepsilon})$  the limit is that of the cubic NLS with local nonlinearity,  $w = \delta$ .

## Higher dimensions

If d > 1 then  $\phi$  has  $\mu_0$ -a.s. negative regularity,  $\phi \notin L^2$ , since  $\sum_{k \in \mathbb{N}} \lambda_k^{-1} = \infty$ . Consequences:

•  $W(\phi) = \frac{1}{2} \int \mathrm{d}x \,\mathrm{d}y \, w(x-y) |\phi(x)|^2 |\phi(y)|^2$  ill-defined even for  $w \in L^\infty$ .

 ${\mbox{ }} p$  -particle correlation functions  $\gamma_p$  are not in trace class, since

$$\operatorname{Tr}(\gamma_1) = \mathbb{E}^{\mu} \|\phi\|_{L^2}^2 = \infty.$$

 On the quantum side, rescaled number of particles N<sub>τ</sub> is no longer bounded. Explicit computation for noninteracting case w = 0:

$$\rho_{\tau}(\mathcal{N}_{\tau}) = \sum_{k \in \mathbb{N}} \frac{1}{\tau} \frac{1}{\mathrm{e}^{\lambda_k/\tau} - 1} \to \infty$$

as  $\tau \to \infty$ . Quantum model has intrinsic cutoff at energies  $\lambda_k \approx \tau$ . Heuristics:

Singularity of classical field  $\iff$  Rapid growth of number of particles .

### Renormalization

Renormalize interaction W by Wick ordering. Formally, take

$$W(\phi) = \frac{1}{2} \int dx \, dy \, w(x-y) (|\phi(x)|^2 - \infty) (|\phi(y)|^2 - \infty) \,.$$

Rigorously, introduce truncated field and density

$$\phi_{[K]} := \sum_{k=0}^{K} \frac{\omega_k}{\sqrt{\lambda_k}} u_k , \qquad \varrho_{[K]}(x) := \mathbb{E}^{\mu_0} |\phi_{[K]}(x)|^2$$

Then

$$W_{[K]} := \frac{1}{2} \int \mathrm{d}x \,\mathrm{d}y \,w(x-y) \big( |\phi_{[K]}(x)|^2 - \varrho_{[K]}(x) \big) \big( |\phi_{[K]}(x)|^2 - \varrho_{[K]}(x) \big)$$

has a limit in  $\bigcap_{p<\infty} L^p(\mu_0)$  as  $K \to \infty$ , denoted by W. Use this W in definition of  $\mu$ . Similarly, we need to renormalize the quantum interaction. The quantum Gibbs state is defined by the renormalized many-body Hamiltonian  $H_{\tau} = H_{\tau,0} + W_{\tau}$ , where

$$W_{\tau} := \frac{1}{2} \int \mathrm{d}x \,\mathrm{d}y \left(\phi_{\tau}^*(x)\phi_{\tau}(x) - \varrho_{\tau}(x)\right) w(x-y) \left(\phi_{\tau}^*(y)\phi_{\tau}(y) - \varrho_{\tau}(y)\right),$$

where the quantum density at  $x \ \varrho_{\tau}(x)$  is defined as

$$\varrho_{\tau}(x) := \rho_{\tau,0} \big( \phi_{\tau}^*(x) \phi_{\tau}(x) \big) \,.$$

Convergence of moments for d = 2, 3

For technical reasons, instead of  $P_{\tau} = e^{-H_{\tau,0}-W_{\tau}}$ , we consider a family of modified thermal quantum states

$$P_{\tau}^{\eta} := e^{-\eta H_{\tau,0}} e^{-(1-2\eta)H_{\tau,0}-W_{\tau}} e^{-\eta H_{\tau,0}}, \qquad \eta \in [0,1).$$

**Theorem** [Fröhlich-K-Schlein-Sohinger, 2016]. Let  $d = 2, 3, w \in L^{\infty}$  positive definite,  $\eta > 0$ , and  $p \in \mathbb{N}$ . Then  $\gamma_{\tau,p}^{\eta} \to \gamma_p$  in Hilbert-Schmidt as  $\tau \to \infty$ . Recent developments:

- [Sohinger, 2019] optimal integrability conditions [Bourgain, 1997] on w:  $w \in L^1$  (d = 1),  $w \in L^{1+}$  (d = 2),  $w \in L^{3+}$  (d = 3).
- [Lewin-Nam-Rougerie, 2018]  $\eta = 0$  for smooth w and d = 2.

#### Counterterm problem

Also works on  $\mathbb{R}^d$  with sufficiently confining potential V. Relation between original and renormalized problems is nontrivial. True Hamiltonian

$$\begin{split} \widetilde{H}_{\tau} &:= \int \mathrm{d}x \,\mathrm{d}y \,\phi_{\tau}^*(x) \big(\nu - \Delta + V\big)(x;y) \phi_{\tau}(y) \\ &\quad + \frac{1}{2} \int \mathrm{d}x \,\mathrm{d}y \,\phi_{\tau}^*(x) \phi_{\tau}^*(y) \,w(x-y) \,\phi_{\tau}(x) \phi_{\tau}(y) \end{split}$$

compared with renormalized Hamiltonian (from above),

$$H_{\tau} = \int \mathrm{d}x \, \phi_{\tau}^*(x) (\kappa - \Delta + v_{\tau}) \phi_{\tau}(x) + \frac{1}{2} \int \mathrm{d}x \, \mathrm{d}y \left( \phi_{\tau}^*(x) \phi_{\tau}(x) - \varrho_{\tau}(x) \right) w(x - y) \left( \phi_{\tau}^*(y) \phi_{\tau}(y) - \varrho_{\tau}(y) \right)$$

Essentially,  $H_{\tau}$  and  $H_{\tau}$  are related by a shift in a diverging chemical potential, provided that one chooses the bare one-body potential  $v_{\tau}$  appropriately (depending on  $\tau$ ).

More precisely, for any constant  $\bar{\varrho}_{\tau} \in \mathbb{R}$  we have

$$\widetilde{H}_{\tau} = H_{\tau} + \left[ \bar{\varrho}_{\tau} \hat{w}(0) - \frac{1}{2\tau} w(0) + \nu - \kappa \right] \mathcal{N}_{\tau} - \frac{1}{2} \int \mathrm{d}x \,\mathrm{d}y \,\varrho_{\tau}(x) w(x-y) \varrho_{\tau}(y) \,,$$

provided that  $v_{\tau}$  solves the counterterm problem

$$v_{\tau} = V + w * \left(\varrho_{\tau}^{v_{\tau}} - \bar{\varrho}_{\tau}\right). \tag{2}$$

For  $\rho_{\tau}^{v_{\tau}} - \bar{\rho}_{\tau}$  to remain bounded, we need  $\lim_{\tau \to \infty} \bar{\rho}_{\tau} = \infty$ , and hence (for bracket to vanish)  $\lim_{\tau \to \infty} \nu = -\infty$ . (Compensates large repulsive interaction energy.)

The counterterm problem (2) is solved in [FKSS, 2016], where we also show that the solution  $v_{\tau}$  converges (in a suitable space) to some v = the correct renormalized external potential.

## Morsels of proof

Basic approach: perturbative expansion of partition functions  $\mathbb{E}^{\mu_0} e^{-zW}$  and  $\operatorname{Tr}(e^{-H_{\tau,0}-zW_{\tau}})$  in powers of z. Well-defined for  $\operatorname{Re} z \ge 0$  but ill-defined for  $\operatorname{Re} z < 0$ : zero radius of convergence around z = 0.

Toy problem: 
$$A(z) = \frac{1}{\sqrt{2\pi}} \int_{\mathbb{R}} \mathrm{d}x \,\mathrm{e}^{-x^2/2} \,\mathrm{e}^{-zx^4}\,;$$

analytic for  ${\rm Re}\,z>0$  but zero radius of convergence, with Taylor coefficient  $a_m=A^{(m)}(0)/m!\sim m!.$ 

However, Taylor series  $\sum_{m \ge 0} a_m z^m$  has Borel transform  $B(z) := \sum_{m \ge 0} \frac{a_m}{m!} z^m$  with positive radius of convergence. Formally, we can recover A from

$$A(z) = \int_0^\infty \mathrm{d}t \,\mathrm{e}^{-t} B(tz) \,.$$

Works provided we can prove good enough bounds on Taylor coefficients and remainder term of A (Sokal, 1980).

Main work: control of the coefficients and remainder of quantum many-body problem. Starting point for algebra is Wick's theorem for the free states.

For time-dependent problem, we perform an expansion of  $\Psi_{\tau}^{t}\Theta_{\tau}(\xi) = e^{it\tau H_{\tau}}\Theta_{\tau}(\xi) e^{-it\tau H_{\tau}}$  in powers of the interaction potential w.

The expansion is controlled graphically, tree graphs sum up precisely to the quantization of  $\Psi^t \Theta(\xi)$ .

Problem: expansion is only convergent on sector of  $\mathcal{F}$  where  $\mathcal{N}_{\tau}$  is bounded. Introduce cutoff in rescaled number of particles  $\mathcal{N}_{\tau}$ . Need to show that for  $f \in C_c^{\infty}(\mathbb{R})$  we have

$$\lim_{\tau \to \infty} \rho_{\tau} \left( \Theta_{\tau}(\xi) f(\mathcal{N}_{\tau}) \right) = \mathbb{E}^{\mu} \left( \Theta(\xi) f(\mathcal{N}) \right)$$
(3)

Problem: cutoff breaks Gaussianity, and Wick's theorem does not apply to (3). Idea: using complex analysis, it suffices to analyse  $\rho_{\tau} (\Theta_{\tau}(\xi) e^{-N_{\tau}})$  for fixed  $\nu > 0$ .