

*Nonlinear diffusion equations as diffusion  
limits of kinetic equations*

Jean DOLBEAULT

*Ceremade, Université Paris Dauphine,  
Place de Lattre de Tassigny,  
75775 Paris Cédex 16, France*

E-mail: [dolbeaul@ceremade.dauphine.fr](mailto:dolbeaul@ceremade.dauphine.fr)

<http://www.ceremade.dauphine.fr/~dolbeaul/>

Będlewo, September 8, 2005

## Outline

- Methods based on self-similarity are now well established at the level of diffusion equations
  - It is an important tool for the understanding of large time asymptotics
  - ⇒ Can we extend such methods to kinetic equations ?
  - Self-similarity: a tool for understanding *dispersive* properties of kinetic equations
  - Notion of *intermediate asymptotics* has to be replaced by *asymptotic stability*
- How to relate kinetic descriptions with diffusive models ?  
*diffusion limits*

## Plan

### I. Time-dependent rescalings and asymptotic stability

A joint work in collaboration with **G. Rein**

### II. Measuring stability in kinetic equations [**J.D.**, **Sánchez**, **Soler**], [**Cáceres**, **Carrillo**, **J.D.**]

### III. Nonlinear diffusion equations as diffusion limits of kinetic equations

A joint work in collaboration with **P. Markowich**, **D. Ölz**, and **C. Schmeiser**. Other contributions by Poupaud, Golse, Goudon, Degond, Schmeiser, Perthame, Ben Abdallah, Chavanis et al., Laurençot, Lemou,...

# I. Time-dependent rescalings and asymptotic stability

Consider the Vlasov-Poisson system

$$\begin{aligned}\partial_t f + v \cdot \nabla_x f - \nabla_x \phi \cdot \nabla_v f &= 0 \\ -\Delta \phi &= \int_{\mathbf{R}^d} f(t, x, v) dv\end{aligned}$$

in the physical space:  $x, v \in \mathbf{R}^3$ , without confinement. Because of the repulsive mean field force, particles **runaway** at infinity and one expects to get **dispersion** estimates at least as  $t \rightarrow \infty$

$$\begin{aligned}\frac{d}{dt} \left( \frac{1}{t} \int \int_{\mathbf{R}^3 \times \mathbf{R}^3} |x - tv|^2 f(t, x, v) dx dv + \frac{1}{2} t \int_{\mathbf{R}^3} |\nabla \phi|^2 dx \right) \\ = -\frac{1}{t^2} \int \int_{\mathbf{R}^3 \times \mathbf{R}^3} |x - tv|^2 f(t, x, v) dx dv \leq 0\end{aligned}$$

Since  $L^p$  norms of the distribution function  $f(t, \cdot)$  are preserved:

$$\|f(t, \cdot)\|_{L^\infty(\mathbf{R}^3 \times \mathbf{R}^3)} \leq \|f_0\|_{L^\infty(\mathbf{R}^3 \times \mathbf{R}^3)} \quad \forall t > 0$$

**Dispersion** can be measured by interpolation

$$0 \leq \rho(t, x) := \int_{\mathbf{R}^d} f(t, x, v) dv = \int_{|x-tv| \leq R} f dv + \int_{|x-tv| > R} f dv$$

$$0 \leq \rho(t, x) \leq 4\pi \left(\frac{R}{t}\right)^3 \|f_0\|_{L^\infty(\mathbf{R}^3 \times \mathbf{R}^3)} + \frac{1}{R^2} \int_{\mathbf{R}^3} |x - tv|^2 f dv$$

Optimizing on  $R = R(x, t)$  with  $x, t$  fixed, we get

$$0 \leq \rho(t, x) \leq C \|f_0\|_{L^\infty(\mathbf{R}^3 \times \mathbf{R}^3)}^{2/5} \left( \int_{\mathbf{R}^3} |x - tv|^2 f dv \right)^{3/5} t^{-6/5}$$

$$\|\rho\|_{L^{5/3}(\mathbf{R}^3)} \leq C \|f_0\|_{L^\infty(\mathbf{R}^3 \times \mathbf{R}^3)}^{2/5} \left( \int \int_{\mathbf{R}^3 \times \mathbf{R}^3} |x - tv|^2 f dx dv \right)^{3/5} t^{-6/5}$$

**Theorem 1.** [Perthame, Illner-Rein] Assume

$$f_0 \in L^1 \cap L^\infty(\mathbb{R}^3 \times \mathbb{R}^3), \int \int_{\mathbb{R}^3 \times \mathbb{R}^3} (|x|^2 + |v|^2) f \, dx \, dv < \infty$$

Then

$$\|\rho(t, \cdot)\|_{L^{5/3}(\mathbb{R}^3)} = O(t^{-3/5})$$

## SCALINGS

If  $(f, \phi)$  is a solution to the Vlasov-Poisson system,

$$f_{\lambda, \mu}(t, x, v) := \lambda^{2-d} f(\lambda t, \mu x, \lambda^{-1} \mu v)$$

and the corresponding  $\phi_{\lambda, \mu}$  given by the Poisson equation are also solutions to the Vlasov-Poisson system.

If we additionally require that the  $L^1(\mathbb{R}^d \times \mathbb{R}^d)$  norm is preserved, we get:  $\mu = \lambda^{2/d}$  and the corresponding distribution function is

$$f_{\lambda, \lambda^{2/d}}(t, x, v) := \lambda^{4-d} f(\lambda t, \lambda^{2/d} x, \lambda^{-1+2/d} v)$$

We immediately see that for  $d = 3$ , this scaling is not convenient since in the singular limit the initial data is not well defined as a measure.

## A SPECIAL SOLUTION

The “monokinetic” distribution function:

$$f_\infty(t, x, v) := \rho(t, x) \delta(v - u(t, x))$$

is a solution of (VP) if

$$\rho(t, x) = \frac{1}{R^d(t)} \mathbb{1}_{B_{R(t)}}(x), \quad u(t, x) = \frac{1}{R(t)} \frac{dR}{dt} x$$

are solutions to the pressureless Euler-Poisson system (EP)

$$\begin{cases} \partial_t \rho + \nabla \cdot (\rho u) = 0 \\ \partial_t u + (u \cdot \nabla_x) u = -\nabla_x \phi \\ -\Delta \phi = \rho \end{cases}$$

$$\iff \frac{dR}{dt} = R^{1-d}$$

## TIME-DEPENDENT RESCALINGS

[Rein, J.D.] We define a change of variables which leaves (VP) as invariant as possible

$$(t, x, v) \mapsto (\tau, \xi, \eta)$$

$$dt = A^2(t)d\tau, \quad x = R(t)\xi$$

Assuming that  $t \mapsto x(t)$  and  $\tau \mapsto \xi(\tau)$  satisfy  $\frac{dx}{dt} = v$  and  $\frac{d\xi}{d\tau} = \eta$   
 $\iff$  *Galilean invariance*, the new velocity variable  $\eta$  has to satisfy

$$v = \frac{dx}{dt} = \dot{R}(t)\xi + R(t)\frac{d\xi}{d\tau}\frac{d\tau}{dt} = \dot{R}(t)\xi + \frac{R(t)}{A^2(t)}\eta$$

Rescaled distribution function:

$$f(t, x, v) = G(t)F(\tau, \xi, \eta)$$

$$d\tau = A^{-2}(t)dt, \quad \xi = R^{-1}(t)x, \quad \eta = \frac{A^2(t)}{R(t)} \left( v - \frac{\dot{R}(t)}{R(t)}x \right).$$

If  $\nu$  and  $W$  are defined as the rescaled spatial density and the rescaled potential respectively, then

$$\nu(\tau, \xi) = \int_{\mathbf{R}^d} F(\tau, \xi, \eta) d\eta = \frac{A^{2d}}{R^d G} \rho(t, x),$$

$$W(\tau, \xi) = \frac{A^{2d}}{R^{d+2} G} \phi(t, x), \quad \nabla_{\xi} W(\tau, \xi) = \frac{A^{2d}}{R^{d+1} G} \nabla_x \phi(t, x),$$

Rescaled Vlasov equation

$$\begin{aligned} \partial_\tau F + \eta \cdot \nabla_\xi F + 2A^2 \left( \frac{\dot{A}}{A} - \frac{\dot{R}}{R} \right) \eta \cdot \nabla_\eta F \\ - \ddot{R} \frac{A^4}{R} \xi \cdot \nabla_\eta F - R^d G A^{4-2d} \nabla_\xi W \cdot \nabla_\eta F + A^2 \frac{\dot{G}}{G} F = 0 \end{aligned}$$

$L^1$  norm is preserved under the change of variables

$$\frac{\dot{A}}{A} - \frac{\dot{R}}{R} = \frac{1}{2d} \frac{\dot{G}}{G} \implies G = \left( \frac{A}{R} \right)^{2d}$$

No time-dependent factor in front of the nonlinear term:

$A = R^{d/4}$ ,  $G = R^{\frac{d-4}{2}d}$  and  $R$  has to solve

$$\ddot{R} + R^{1-d} = 0$$

Rescaled Vlasov-Poisson system (RVP):

$$\begin{aligned} \partial_\tau F + \eta \cdot \nabla_\xi F + \nabla_\eta \left[ \left( \xi - \nabla_\xi W + \frac{d-4}{2} R^{\frac{d}{2}-1} \dot{R} \eta \right) F \right] &= 0 \\ -\Delta W = \nu(\tau, \xi) &= \int_{\mathbf{R}^d} F(\tau, \xi, \eta) d\eta \end{aligned}$$

The relation between the old and the new variables is

$$\begin{aligned} dt &= R^{d/2} d\tau, & d\tau &= R^{-d/2} dt \\ x &= R\xi, & \xi &= R^{-1}x \\ v &= \dot{R}\xi + R^{1-\frac{d}{2}}\eta, & \eta &= R^{\frac{d}{2}-1} \left( v - \frac{\dot{R}}{R}x \right) \end{aligned}$$

and the rescaled functions are given by

$$\begin{aligned} F(\tau, \xi, \eta) &= R^{\frac{4-d}{2}d} f(t, x, v), & \nu(\tau, \xi) &= R^d \rho(t, x) \\ W(\tau, \xi) &= R^{d-2} \phi(t, x), & \nabla_\xi W(\tau, \xi) &= R^{d-1} \nabla_x \phi(t, x) \end{aligned}$$

$F_\infty^M(\xi, \eta) = \nu_\infty^M(\xi)\delta(\eta)$  where  $\delta$  is the usual Dirac distribution is a steady state, with

$$\nabla_\xi W_\infty^M(\xi) = \begin{cases} \xi & \text{if } |\xi| \leq (M/|S^{d-1}|)^{1/d} \\ \xi/|\xi|^d & \text{if } |\xi| > (M/|S^{d-1}|)^{1/d} \end{cases}$$

and associated spatial density

$$\nu_\infty^M(\xi) = d \cdot \mathbb{I}_{B^d((M/|S^{d-1}|)^{1/d})}$$

By the inverse rescaling transformation, we get

$$f_\infty^M(t, x, v) = \frac{d}{R(t)^d} \mathbb{I}_{B^d(R(t)(M/|S^{d-1}|)^{1/d})}(x) \delta\left(v - \frac{\dot{R}(t)}{R(t)}x\right)$$

and  $\rho_\infty^M(t, x) = \frac{d}{R(t)^d} \mathbb{I}_{B^d(R(t)(M/|S^{d-1}|)^{1/d})}(x)$ ,  $u_\infty^M(t, x) = \frac{\dot{R}(t)}{R(t)}x$

This defines a weak solution of (VP) or (EP)

The behavior of  $R(t)$  depends on the dimension. For  $d \neq 2$ ,

$$\frac{1}{2} \left( \frac{dR}{dt}(t) \right)^2 - \frac{1}{2} \left( \frac{dR}{dt}(0) \right)^2 = \frac{1}{2-d} R^{2-d}(t) - R^{2-d}(0) = 0$$

and, for  $d = 2$ ,

$$\frac{1}{2} \left( \frac{dR}{dt}(t) \right)^2 - \frac{1}{2} \left( \frac{dR}{dt}(0) \right)^2 = \log \left( \frac{R(t)}{R(0)} \right)$$

*If we choose  $\frac{dR}{dt}(0) = 0$  and  $R(0) = 1$ , initial data for the rescaled problem is the same as for the original system*

As  $t \rightarrow \infty$ , we get the following equivalences

$$\begin{array}{lll} R(t) \sim t^2 & \text{if} & d = 1 \\ R(t) \sim t \sqrt{\log t} & \text{if} & d = 2 \\ R(t) \sim t & \text{if} & d \geq 3 \end{array}$$

In terms of the rescaled time variable, this means

$$\begin{aligned} \tau(t) &\sim \log t && \text{if } d = 1 \\ \tau(t) &\sim \sqrt{\log t} && \text{if } d = 2 \\ \tau(t) &\rightarrow \tau_\infty < \infty && \text{if } d \geq 3 \end{aligned}$$

The friction coefficient

$$\beta := \frac{d-4}{2} R^{d-1} \dot{R}$$

is not constant, but converges as  $t \rightarrow \infty$  to a positive constant, at least for  $d < 4$ . In the case  $d = 4$ , one recovers that the Vlasov-Poisson system is conformally invariant

Decay of the energy of the rescaled system

$$\begin{aligned} \frac{d}{d\tau} \left[ \int \int_{\mathbf{R}^d \times \mathbf{R}^d} \left( \frac{1}{2} |\eta|^2 + \frac{1}{2} |\xi|^2 + \frac{1}{2} W \right) F(\tau, \xi, \eta) d\xi d\eta \right] \\ = -\beta(\tau) \int \int_{\mathbf{R}^d \times \mathbf{R}^d} |\eta|^2 F(\tau, \xi, \eta) d\xi d\eta \leq 0 \end{aligned}$$

## II. Measuring stability in kinetic equations

Consider the (VP) system in the gravitational case

$$\begin{cases} \partial_t f + v \cdot \nabla_x f - (\nabla_x \phi + \nabla_x \phi_0) \cdot \nabla_v f = 0 \\ \Delta_x \phi = 4\pi\rho, \quad \lim_{|x| \rightarrow \infty} \phi(t, x) = 0 \end{cases} \quad (VP)$$

in presence of an external potential  $\phi_0$ . If we look for stationary solutions taking the form

$$f(x, v) = \gamma \left( \frac{1}{2} |v|^2 + \phi(x) + \phi_0(x) - \kappa \right)$$

Vlasov's equation is satisfied and the problem is reduced to solve the nonlinear Poisson equation

$$\Delta \phi = 4\pi G(\phi + \phi_0 - \kappa)$$

with

$$G(u) := \int_{\mathbf{R}^3} \gamma \left( \frac{1}{2} |v|^2 + u \right) dv = \frac{4\pi}{3} \int_0^{+\infty} \gamma(s + u) \sqrt{2s} ds$$

## A POTENTIAL ENERGY ESTIMATE

**Lemma 2.** *There exists a positive constant  $C$  such that*

$\forall f \in L^1_{+} \cap L^{\infty}(\mathbb{R}^6)$  with  $|v|^2 f \in L^1(\mathbb{R}^6)$

$$\int_{\mathbb{R}^3} |\nabla \phi|^2 dx \leq C \|f\|_{L^1(\mathbb{R}^6)}^{7/6} \|f\|_{L^{\infty}(\mathbb{R}^6)}^{1/3} \left( \int_{\mathbb{R}^6} |v|^2 f(x, v) dx dv \right)^{1/2}$$

$$\int_{\mathbb{R}^3} |\nabla \phi|^2 dx = \int_{\mathbb{R}^3} (-\Delta \phi) \phi dx = 4\pi \int_{\mathbb{R}^6} \frac{\rho(y)\rho(x)}{|x-y|} dx dy$$

According to the **Hardy-Littlewood-Sobolev** inequalities,

$$\int_{\mathbb{R}^3} |\nabla \phi|^2 dx \leq 4\pi \Sigma \|\rho\|_{L^{6/5}(\mathbb{R}^3)}^2$$

Because of **Hölder's** inequality,  $\|\rho\|_{L^{6/5}(\mathbb{R}^3)} \leq \|\rho\|_{L^1(\mathbb{R}^3)}^{7/12} \|\rho\|_{L^{5/3}(\mathbb{R}^3)}^{5/12}$

Use **interpolation inequalities** to bound  $\|\rho\|_{L^{5/3}(\mathbb{R}^3)}$

## AN EQUIVALENT MINIMIZATION PROBLEM

$\Gamma_M = \{f \in L^1 \cap L^\infty(\mathbb{R}^6) : f(x, v) \geq 0, \|f\|_{L^1(\mathbb{R}^6)} = M, \|f\|_{L^\infty(\mathbb{R}^6)} \leq 1\}$   
Let  $J_M = \inf \{J(f) : f \in \Gamma_M\}$ ,

$$J(f) = \frac{\frac{1}{2} \int_{\mathbb{R}^6} |v|^2 f \, dx \, dv}{\left(\frac{1}{8\pi} \int_{\mathbb{R}^6} |\nabla \phi|^2 \, dx\right)^2} \equiv \frac{E_{KIN}(f)}{(E_{POT}(f))^2}$$

**Lemma 3.** *The minimization problems  $E(f) = E_M$  and  $J(f) = J_M$  over the set  $\Gamma_M$  are equivalent*

(i) *Their respective minima satisfy*

$$4 J_M E_M = -1$$

(ii) *If  $f_M \in \Gamma_M$  is a minimizer of the functional  $E$ , then it is also a minimizer of the functional  $J$*

**Theorem 1** (J.D., Sánchez, Soler). *Let  $f_M$  be a minimizing function for the functional  $E$  on  $\Gamma_M$ , with radial mass density. Then*

$$f_M(x, v) = \begin{cases} 1 & \text{if } \frac{1}{2}|v|^2 + \phi_{f_M}(x) < \frac{7}{3} \frac{E(f_M)}{M} \\ 0 & \text{otherwise} \end{cases}$$

where  $\phi_{f_M}$  is the unique radial solution on  $\mathbb{R}^3$  of

$$\Delta \phi_{f_M} = \frac{1}{3} (4\pi)^2 \left[ 2 \left( \frac{7}{3} \frac{E_M}{M} - \phi_{f_M} \right)_+ \right]^{3/2}$$

*It is the unique minimizer with radial mass density and it is also a steady-state solution to the VP system. Moreover, if  $f$  is another minimizing function, then with  $\bar{x} := \frac{1}{M} \int_{\mathbb{R}^6} x f(x, v) dx dv$*

$$f(x, v) = f_M(x - \bar{x}, v) \quad \forall (x, v) \in \mathbb{R}^6$$

## NONLINEAR STABILITY FOR THE EVOLUTION PROBLEM

[Guo, Rein, Wolansky, Sánchez, Soler, Schaeffer, Lemou-Méhats-Raphaël]  
Consider for any  $g, h \in \Gamma_M$  the distance  $d$  defined by

$$d(g, h) = E(g) - E(h) + \frac{1}{4\pi} \|\nabla\phi_g - \nabla\phi_h\|_{L^2(\mathbf{R}^3)}^2$$

**Theorem 2.** *For every  $\epsilon > 0$ , there exists a  $\delta > 0$  such that, if  $f$  is a solution of the VP system with an initial condition  $f_0 \in \Gamma_M$ , then*

$$d(f_0, f_M) \leq \delta \implies d(f^*(t), f_M) \leq \epsilon \quad \forall t \geq 0$$

The result is easily achieved by contradiction since  $E(f^*(t)) - E(f_M) \leq E(f_0) - E(f_M) \searrow 0$  implies  $\|\nabla\phi_{f^*(t)} - \nabla\phi_{f_M}\|_{L^2(\mathbf{R}^3)} \searrow 0$   
 $\square$

### III. Nonlinear diffusion equations as diffusion limits of kinetic equations

## MODEL

$$\epsilon^2 \partial_t f + \epsilon v \cdot \nabla_x f - \epsilon \nabla_x V(x) \cdot \nabla_v f = Q[f] \quad (1)$$

$$Q[f] := G_f - f$$

$$G_f := \gamma \left( \frac{1}{2} |v|^2 + V(x) - \mu_{\rho_f}(x, t) \right)$$

Local Fermi level:  $\mu_{\rho_f}$  is implicitly determined by the condition

$$\int_{\mathbb{R}^3} G_f dv = \rho_f := \int_{\mathbb{R}^3} f dv$$

The collision operator can be rewritten as

$$\mu_{\rho_f}(x, t) = V(x) + \bar{\mu}(\rho_f(x, t)) \quad \text{and} \quad \int_{\mathbb{R}^3} \gamma \left( \frac{1}{2} |v|^2 - \bar{\mu}(\rho_f) \right) dv = \rho_f$$

$$Q[f] = G_f - f, \quad G_f = \gamma \left( \frac{1}{2} |v|^2 - \bar{\mu}(\rho_f) \right)$$

## ASSUMPTIONS

The energy profile  $\gamma : (E_1, E_2) \rightarrow \mathbb{R}_+$  is a nonincreasing, non-negative  $\mathcal{C}^1$  function,  $-\infty \leq E_1 < E_2 \leq \infty$ ,  $\lim_{E \rightarrow E_2} \gamma(E) = 0$ . If  $E_2 < \infty$ , we extend  $\gamma$  to  $[E_2, \infty)$  by 0 and assume that there are constants  $k > 0$  and  $C > 0$  such that

$$\gamma(E) \leq C(E_2 - E)^k \quad \text{on} \quad (\hat{E}, E_2)$$

If  $E_2 = \infty$  we require

$$\gamma(E) = O(E^{-5/2}) \quad \text{as} \quad E \rightarrow \infty$$

to ensure existence of second velocity moments

+ technical assumptions on  $\gamma$  close to  $E_2$

## FORMAL ASYMPTOTIC AS $\epsilon \rightarrow 0$

$$f = \sum_{i=0}^{\infty} f^i \epsilon^i \quad \rho^i := \int_{\mathbb{R}^3} f^i dv \quad \rho = \sum_{i=1}^{\infty} \rho^i \epsilon^i$$

Let  $G^i := \text{sign}(\rho^i) \gamma(|v|^2/2 - \bar{\mu}(|\rho^i|))$ ,  $G_f \approx \sum_{i=1}^{\infty} G^i \epsilon^i$

$$\epsilon^0 : \quad G^0(x, v, t) = G_{f^0} = \gamma(|v|^2/2 + V(x) - \mu^0(x, t)) = f^0$$

$$\epsilon^1 : \quad v \cdot \nabla_x f^0 - \nabla_x V \cdot \nabla_v f^0 = G^1 - f^1$$

$$\epsilon^2 : \quad \partial_t f^0 + v \cdot \nabla_x f^1 - \nabla_x V \cdot \nabla_v f^1 = G^2 - f^2$$

$$\partial_t \int_{\mathbb{R}^3} f^0 dv + \nabla_x \cdot \int_{\mathbb{R}^3} v f^1 dv = O(\epsilon)$$

$$f^1 = v \cdot \nabla_x \mu^0 \gamma' \left( \frac{1}{2} v^2 + V(x) - \mu^0(x, t) \right) + G^1 \quad \int_{\mathbb{R}^3} v f^1 dv = -\rho^0 \nabla_x \mu^0$$

Collecting these estimates, we get, for  $\rho^0(x, t) = \int_{\mathbb{R}^3} f^0(x, v, t) dv$ ,

$$\partial_t \rho^0 = \nabla \cdot (\rho^0 \nabla \mu^0)$$

Use:  $\mu^0 = \bar{\mu}(\rho^0) + V$  to recover the expected drift-diffusion equation :

$$\partial_t \rho^0 = \nabla \cdot (\rho^0 (\nabla \bar{\mu}(\rho^0) + \nabla V(x))) \quad (2)$$

## MOTIVATION

- collisions : short time scale
- Gibbs states are usually better known than collision kernels
- Gibbs states  $\iff$  generalized entropies
- nonlinear diffusion equations are difficult to justify directly
- global Gibbs states are the same at the kinetic / diffusion levels

[Ben Abdallah, J.D.], [Chavanis, Laurençot, Lemou], [Degond, Ringhofer]

**Example 1.** Power law case :  $\gamma(E) := DE^{-k}$ ,  $D > 0$ ,  $k > 5/2$   
 (existence of second velocity moments)

$$\bar{\mu}(\rho) = -\left(\frac{\rho}{D\beta(k)}\right)^{\frac{1}{\frac{3}{2}-k}}, \quad \text{where} \quad \beta(k) := 4\pi\sqrt{2} \int_0^\infty \frac{\sqrt{s}}{(s+1)^k} ds$$

*Fast diffusion equations :*

$$\partial_t \rho = \nabla \cdot \left( \Theta \nabla \left( \rho^{\frac{k-5/2}{k-3/2}} \right) + \rho \nabla V \right), \quad \text{where} \quad \Theta := \frac{1}{k - \frac{5}{2}} \left( \frac{1}{D\beta(k)} \right)^{\frac{1}{\frac{3}{2}-k}}$$

Outside of a finite ball, the potential grows faster than a power

$$V(x) \geq C|x|^q, \quad \text{a.e. for } |x| > R \quad \text{with} \quad q > \frac{3}{k - 5/2}.$$

**Example 2.** Maxwell distribution :  $\gamma(E) = \exp(-E)$

$$\bar{\mu}(\rho) = \log \rho - \frac{3}{2} \log(2\pi)$$

Linear drift-diffusion equation :  $\nu(\rho) = \rho$

$$\partial_t \rho = \nabla \cdot (\nabla \rho + \rho \nabla V)$$

*"the linear case"*

Growth assumption on the potential

$$V(x) \geq q \log(|x|), \quad \text{a.e. for } |x| > R \quad \text{with } q > 3$$

**Example 3.** Let  $\gamma$  be a cut-off power with positive exponent :

$$\gamma(E) = (E_2 - E)_+^k := \begin{cases} D(E_2 - E)^k & \text{for } E < E_2, \quad D > 0, \quad k > 0 \\ 0 & \text{otherwise} \end{cases}$$

$$\bar{\mu}(\rho) = \left( \frac{\rho}{D\alpha(k)} \right)^{\frac{1}{k+\frac{3}{2}}} - E_2, \quad \text{where } \alpha(k) = 4\pi\sqrt{2} \int_0^1 \sqrt{u}(1-u)^k du$$

*Porous medium equations :*  $\nu(\rho) = \Theta \rho^{\frac{2k+5}{2k+3}}$

$$\partial_t \rho = \nabla \cdot \left( \Theta \nabla \left( \rho^{\frac{k+5/2}{k+3/2}} \right) + \rho \nabla V \right), \quad \text{where } \Theta := \frac{1}{k + \frac{5}{2}} \left( \frac{1}{D\alpha(k)} \right)^{\frac{1}{k+\frac{3}{2}}}$$

Growth condition on the potential : if  $\mu^*$  is the upper bound for the Fermi energy

$$\left( E_2 + \mu^* - V(x) \right)_+ = O\left( \frac{1}{|x|^q} \right) \quad \text{a.e. as } |x| \rightarrow \infty, \quad q > \frac{3}{k + \frac{3}{2}}$$

**Example 4.** Fermi-Dirac distribution :  $\gamma(E) = \frac{1}{\exp(E) + \alpha}$

$$(\bar{\mu}^{-1})(\theta) = \frac{4\pi\sqrt{2}}{\alpha} \int_0^\infty \frac{\sqrt{p} dp}{\exp(p - \theta - \log \alpha) + 1} = -\frac{(2\pi)^{\frac{3}{2}}}{\alpha} \text{Li}_{3/2}(-\alpha \exp(\theta))$$

$$\bar{\mu}(\rho) = \log \left( -\frac{1}{\alpha} \left( \text{Li}_{3/2}^{-1} \right) \left( -\frac{\alpha \rho}{(2\pi)^{3/2}} \right) \right), \quad \text{Li}_n(z) := \sum_{k=1}^{\infty} \frac{z^k}{k^n}$$

Macroscopic equation :  $\partial_t \rho = \nabla \cdot \left( (D(\rho) \nabla \rho + \rho \nabla V) \right)$

$$D(\rho) = \nu'(\rho) = \rho \bar{\mu}'(\rho) = \frac{-\alpha}{(2\pi)^{3/2}} \frac{\rho}{\text{Li}_{1/2} \left( \left( \text{Li}_{3/2}^{-1} \right) \left( \frac{-\alpha \rho}{(2\pi)^{3/2}} \right) \right)}$$

Moreover the expansion of  $D(\rho)$  at  $\rho = 0$  gives

$$D(\rho) = 1 + \frac{\sqrt{2}}{4} \frac{\alpha \rho}{(2\pi)^{3/2}} + \left( \frac{3}{8} - \frac{2\sqrt{3}}{9} \right) \frac{\alpha^2 \rho^2}{(2\pi)^3} + O(\rho^3)$$

**Example 5.** Bose-Einstein distribution :  $\gamma(E) = \frac{1}{\exp(E) - \alpha}$

$$(\bar{\mu}^{-1})(\theta) = \frac{4\pi\sqrt{2}}{\alpha} \int_0^\infty \frac{\sqrt{p} dp}{\exp(p - \theta - \log \alpha) - 1} = \frac{(2\pi)^{\frac{3}{2}}}{\alpha} \text{Li}_{3/2}(\alpha \exp(\theta))$$

$$\bar{\mu}(\rho) = \log \left( \frac{1}{\alpha} \left( \text{Li}_{3/2}^{-1} \right) \left( \frac{\alpha \rho}{(2\pi)^{3/2}} \right) \right)$$

Macroscopic equation :  $\partial_t \rho = \nabla \cdot \left( (D(\rho) \nabla \rho + \rho \nabla V) \right)$

$$D(\rho) = \nu'(\rho) = \rho \bar{\mu}'(\rho) = \frac{\alpha}{(2\pi)^{3/2}} \frac{\rho}{\text{Li}_{1/2} \left( \left( \text{Li}_{3/2}^{-1} \right) \left( \frac{\alpha \rho}{(2\pi)^{3/2}} \right) \right)}$$

Observe that  $\lim_{\rho \rightarrow \bar{\rho}} \nu'(\rho) = 0$  and  $\lim_{\rho \rightarrow 0} \nu'(\rho) = 1$

Maximal density  $\bar{\rho}$  :  $\bar{\rho} = \frac{(2\pi)^{3/2} \zeta(\frac{3}{2})}{\alpha} \approx \frac{41,144}{\alpha}$

Local Gibbs state and diffusion coefficient : *Modelization*

$$Q[f] = G_f - f, \quad G_f = \gamma \left( \frac{1}{2} |v|^2 - \bar{\mu}(\rho_f) \right)$$

$\bar{\mu}^{-1} : (-E_2, -E_1) \rightarrow (0, \infty)$  is such that

$$(\bar{\mu}^{-1})(\theta) = 4\pi\sqrt{2} \int_0^\infty \gamma(p - \theta) \sqrt{p} dp$$

We extend  $\bar{\mu}^{-1}$  by the value 0 on  $(-\infty, -E_2)$ . Differentiation with respect to  $\theta$  leads to the Abelian equation

$$\frac{(\bar{\mu}^{-1})'(\theta)}{2\pi\sqrt{2}} = \int_{-\infty}^\theta \frac{\gamma(-q)}{\sqrt{\theta - q}} dq$$

and gives an explicit expression of  $\gamma$  in terms of  $\bar{\mu}^{-1}$

$$\gamma(E) = \frac{1}{\sqrt{2}} \frac{d^2}{2\pi^2 dE^2} \int_{-\infty}^{-E} \frac{(\bar{\mu}^{-1})(\theta)}{\sqrt{-E - \theta}} d\theta$$

## INITIAL DATA

We assume that there is a constant *Fermi level*  $\mu^*$  such that

$$0 \leq f_I(x, v) \leq f^*(x, v) := \gamma\left(\frac{1}{2}|v|^2 + V(x) - \mu^*\right) \quad \forall (x, v) \in \mathbb{R}^6$$

Maximal macroscopic density :

$$\bar{\rho} := \lim_{\theta \rightarrow -E_1^+} \int_{\mathbb{R}^3} \gamma\left(\frac{1}{2}|v|^2 - \theta\right) dv$$

If  $\bar{\rho} < \infty$  we assume

$$\bar{\mu}^{-1}(\mu^*) \leq \bar{\rho}$$

## POTENTIAL

$\nabla_x V \in W^{1,\infty}\mathbb{R}^3$ ) and  $V$  is bounded from below

$$\inf_{x \in \mathbb{R}^3} V(x) = V_{\min} = 0$$

*Confinement condition* : with  $f^*(x, v) := \gamma\left(\frac{1}{2}|v|^2 + V(x) - \mu^*\right)$

$$f^* \in L^1(\mathbb{R}^3 \times \mathbb{R}^3) \quad \text{and} \quad \iint_{\mathbb{R}^6} \left(\frac{1}{2}|v|^2 + V(x)\right) f^*(x, v) dv dx < \infty$$

Observe that this implies  $\|f^*\|_{L^1} \geq \|f_I\|_{L^1} = M$

A *compatibility assumption* : Given a Gibbs state (a function  $\gamma$ ), impose some minimal growth conditions on  $V$ .

## EXISTENCE AND UNIQUENESS

**Proposition 4.** *For any  $p \in (1, \infty)$ , Eq. (1) has a unique weak solution in*

$$\mathcal{V} := \{f \in \mathcal{C}(0, \infty; (L^1 \cap L^p)(\mathbb{R}^6)) : 0 \leq f \leq f^*, \forall t > 0 \text{ a.e.}\}$$

Proof: Cf. [Poupaud-Schmeiser, 1991] : define the map  $f \mapsto \Gamma[f] = g$

$$\begin{aligned} \epsilon^2 \partial_t g + \epsilon v \cdot \nabla_x g - \epsilon \nabla_x V \cdot \nabla_v g &= G_f - g \\ g(t=0) &= f_I \end{aligned}$$

$\Gamma$  maps  $\mathcal{V}$  into itself and is a contraction for sufficiently small time intervals.

## FREE ENERGY

$$\mathcal{F}[f] := \iint_{\mathbb{R}^6} \left[ \left( \frac{1}{2}|v|^2 + V \right) f + \beta_\gamma[f] \right] dv dx$$

$$\beta_\gamma[f] := \int_0^f -\gamma^{-1}(s) ds$$

$-\gamma^{-1}$  is monotonically increasing  $\implies \beta_\gamma$  is a convex function  
Microscopic energy associated to a distribution function  $f$  :

$$E_f(x, v, t) := \frac{1}{2}|v|^2 + V(x) - \mu_{\rho_f}(x, t) = \frac{1}{2}|v|^2 - \bar{\mu}(\rho_f(x, t)) = -\gamma^{-1}[G_f]$$

$$\epsilon^2 \frac{d}{dt} \mathcal{F}(f(\cdot, \cdot, t)) = \iint_{\mathbb{R}^6} (G_f - f) (\gamma^{-1}[G_f] - \gamma^{-1}[f]) dv dx := -D[f] \leq 0$$

## CONSEQUENCES

$$\epsilon^2 \left[ \mathcal{F}(f(\cdot, \cdot, \cdot, t)) - \mathcal{F}(f_I) \right] = - \int_0^T D[f](t) dt$$

$$\mathcal{F}_{\text{loc}}[f](x, \cdot, t) := \int_{\mathbb{R}^3} \left[ \left( \frac{1}{2}|v|^2 + V(x) - \mu_f(x, t) \right) f(x, v, t) + \beta_\gamma(f(x, v, t)) \right] dv$$

is convex. Minimum if and only if  $f = G_f$

$$0 = \frac{1}{2}|v|^2 + V(x) - \mu_f(x, t) + \beta'_\gamma[f] = \frac{1}{2}|v|^2 + V(x) - \mu_f(x, t) - \gamma^{-1}[f]$$

$$\mathcal{F}_{\text{loc}}[f](x, t) \geq \mathcal{F}_{\text{loc}}[G_f](x, t) \quad \text{and} \quad \mathcal{F} \geq \mathcal{F}[G_f] \geq \mathcal{F}[g]$$

$$\mathcal{F}[g] = \iint_{\mathbb{R}^6} \gamma \left( \frac{1}{2}|v|^2 + V - \mu \right) \left( \mu - \frac{|v|^2}{3} \right) dv dx = \int_{\mathbb{R}^3} \left( \mu \rho_g - \nu(\rho_g) \right) dx$$

$$\text{with} \quad \rho_g = \bar{\mu}^{-1}(\mu - V) \quad \nu(\rho_g) := \frac{1}{3} \int_{\mathbb{R}^3} |v|^2 \gamma \left( \frac{1}{2}|v|^2 - \bar{\mu}(\rho_g) \right) dv$$

Consider the partition of the support of  $f^*$  according to

$$\Omega_+ := \left\{ (x, v, t) \in \text{supp } f^* \subset \mathbb{R}^6 \times (0, T) : E_f = \frac{1}{2}|v|^2 - \bar{\mu}(\rho_f(x, t)) < E_2 \right\}$$

$$\Omega_0 := \left\{ (x, v, t) \in \text{supp } f^* \subset \mathbb{R}^6 \times (0, T) : E_f = \frac{1}{2}|v|^2 - \bar{\mu}(\rho_f(x, t)) \geq E_2 \right\}$$

*The family of solutions  $f^\epsilon$ , up to the extraction of a subsequence, converges to its local Gibbs state  $G_{f^\epsilon} = \gamma(E_{f^\epsilon})$  a.e. on  $\Omega_+$  and it converges to 0 a.e. on  $\Omega_0$ .*

$$\Omega_+^{x,t} := \{v \in \mathbb{R}^3 : (x, v, t) \in \Omega_+\} \quad \text{and} \quad \Omega_0^{x,t} := \{v \in \mathbb{R}^3 : (x, v, t) \in \Omega_0\}$$

Notice that  $\Omega_+^{x,t} = \mathbb{R}^3$  if  $E_2 = \infty$

**Lemma 5.** *For any nonnegative function  $f \leq f^*$  there exists a constant, which does not depend on  $x$  and  $t$ , such that*

$$\int_{\Omega_+^{x,t}} v_i^{2m} \frac{G_f - f}{\gamma^{-1}[f] - E_f} dv \leq \mathcal{M},$$

for any  $m = 1, 2$ ,  $i = 1, 2, 3$ .

Scaled perturbations of the first and second moments

$$j^\epsilon := \int_{\mathbb{R}^3} v \frac{f^\epsilon - G f^\epsilon}{\epsilon} dv \quad \text{and} \quad \kappa^\epsilon := \int_{\mathbb{R}^3} v \otimes v \frac{f^\epsilon - G f^\epsilon}{\epsilon} dv$$

**Lemma 6.** For any bounded, open set  $U \subset \mathbb{R}^3 \times [0, T)$ , there are two constants  $\mathcal{M}_U^1$  and  $\mathcal{M}_U^2$ , which do not depend on  $\epsilon$ , such that

$$\|j^\epsilon\|_{L^2_{x,t}(U)} \leq \mathcal{M}_U^1 \quad \text{and} \quad \|\kappa^\epsilon\|_{L^2_{x,t}(U)} \leq \mathcal{M}_U^2 \quad \text{as} \quad \epsilon \rightarrow 0.$$

If  $g(x, v, t) := \gamma(|v|^2/2 - \bar{\mu}(\rho(x, t)))$ , then

$$\int_{\mathbb{R}^3} v \otimes v g dv = \nu(\rho) \text{Id}^{3 \times 3} \quad \text{where} \quad \nu(\rho) := \int_0^\rho \sigma \bar{\mu}'(\sigma) d\sigma$$

**Proposition 7.**  $\rho^\epsilon \rightarrow \rho^0$  in  $L^p_{loc}$  strongly for all  $p \in (1, \infty)$ .

Div-Curl Lemma as in [Goudon-Poupaud, 2001]. Integrate (1) with respect to  $dv$  and  $v dv$

$$\begin{cases} \partial_t \rho^\epsilon + \nabla_x \cdot j^\epsilon = 0 \\ \epsilon^2 \partial_t j^\epsilon + \nabla_x \cdot \int_{\mathbb{R}^3} v \otimes v f^\epsilon dv = -j^\epsilon - \rho^\epsilon \nabla_x V \end{cases}$$

Split the second moments of  $f$

$$\int_{\mathbb{R}^3} v \otimes v f^\epsilon dv = \int_{\mathbb{R}^3} v \otimes v G^{f^\epsilon} dv + \int_{\mathbb{R}^3} v \otimes v (f^\epsilon - G^{f^\epsilon}) dv = \nu(\rho^\epsilon) I^{3 \times 3} + \epsilon \kappa^\epsilon$$

$$\begin{cases} \partial_t \rho^\epsilon + \nabla_x \cdot j^\epsilon = 0 \\ \nabla_x \nu(\rho^\epsilon) = -j^\epsilon - \rho^\epsilon \nabla_x V - \epsilon \nabla_x \cdot \kappa^\epsilon - \epsilon^2 \partial_t j^\epsilon \end{cases}$$

Apply the Div-Curl Lemma to

$$U^\epsilon := (\rho^\epsilon, j^\epsilon), \quad V^\epsilon := (\nu(\rho^\epsilon), 0, 0, 0)$$

With  $(\text{curl } w)_{ij} = w_{x_j}^i - w_{x_i}^j$  and

$$\begin{cases} \text{div}_{t,x} U^\epsilon = 0, \\ (\text{curl}_{t,x} V^\epsilon)_{1,2\dots 4} = -j^\epsilon - \rho^\epsilon \nabla_x V - \epsilon \nabla_x \cdot \kappa^\epsilon - \epsilon^2 \partial_t j^\epsilon \end{cases}$$

we obtain the convergence of  $U^{\epsilon_i} \cdot V^{\epsilon_i} = \rho^{\epsilon_i} \nu(\rho^{\epsilon_i})$

As in [Marcati-Milani, 1990], we deduce using Young measures that the convergence of  $\rho^{\epsilon_i}$  is strong. The strict convexity assumption is replaced by the strict monotonicity of the function  $\nu$  in  $\rho \nu(\rho)$ .

**Theorem 8.** For any  $\varepsilon > 0$ , the equation has a unique weak solution  $f^\varepsilon \in C(0, \infty; L^1 \cap L^p(\mathbb{R}^6))$  for all  $p < \infty$ . As  $\varepsilon \rightarrow 0$ ,  $f^\varepsilon$  weakly converges to a local Gibbs state  $f^0$  given by

$$f^0(x, v, t) = \gamma \left( \frac{1}{2} |v|^2 + V(x) - \bar{\mu}(\rho(x, t)) \right) \quad \forall (x, v, t) \in \mathbb{R}^3 \times \mathbb{R}^3 \times \mathbb{R}_+$$

where  $\rho$  is a solution of the nonlinear diffusion equation

$$\partial_t \rho = \nabla_x \cdot (\nabla_x \nu(\rho) + \rho \nabla_x V(x))$$

with initial data  $\rho(x, 0) = \rho_I(x) := \int_{\mathbb{R}^3} f_I(x, v) dv$

$$\nu(\rho) = \int_0^\rho s \bar{\mu}'(s) ds$$

Moreover,  $\int_{\mathbb{R}^3} f^\varepsilon dv$  strongly converges to  $\rho$  in  $L^p_{loc}$  as  $\varepsilon \rightarrow 0$