

CREEP TIDE THEORY

**Synchronization,
Dissipation ,
Circularization**

**S.Ferraz-Mello (USP)
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G.H.Darwin, 1880

Goldreich, 1963

Kaula, 1964

Alexander, 1973

Zahn, 1977

Mignard, 1979

Hut, 1981

Eggleton &c, 1998

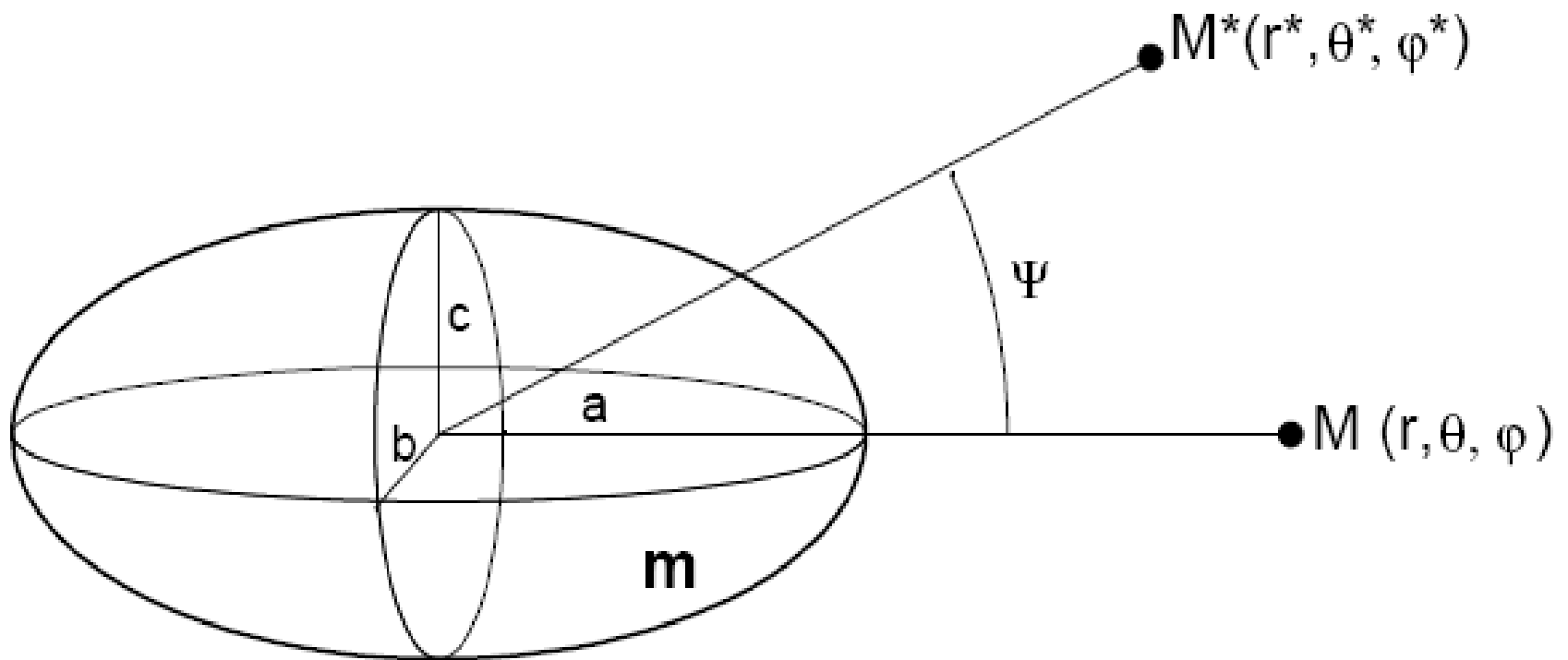
Mardling & Lin, 2002

Ferraz-Mello &c, 2008 (**REVIEW**)

and many others

Static tide (Jeans prolate spheroid)

$$\epsilon = \frac{a}{b} - 1 = \frac{15}{4} \left(\frac{M}{m} \right) \left(\frac{R}{r} \right)^3.$$



Potential at M*

$$\mathbf{U} = -\frac{Gm}{r^*} - \frac{k_f GM R^5}{2r^{*3}r^{*+3}}(3 \cos^2 \Psi - 1);$$

$$k_f = 15A/4mR^2 \quad (\text{Fluid Love number})$$

Dynamical tide

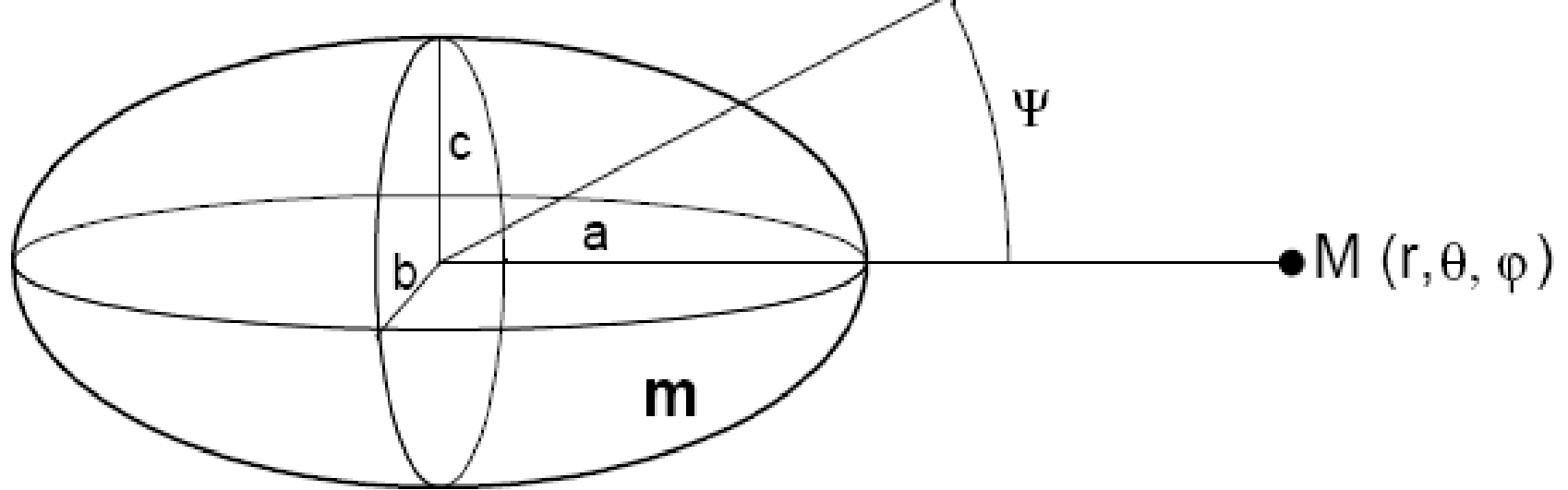
Introduces phase lags:

$$\cos \psi \rightarrow \cos (\psi - \varepsilon)$$

Decomposes :

$$\cos (\psi - \varepsilon) \Rightarrow \cos \psi + \varepsilon \sin \psi$$

elastic+anelastic



Force acting on M (due to **elastic tide**)

$$F_1 = \mathcal{F}(r)$$

$$F_2 = 0$$

$$F_3 = 0$$



$$\text{torque} = 0$$

$$\langle W \rangle = \text{dissipation} = 0$$

$$de/dt = 0$$

Recent theories (of the anelastic tide):

Efroimsky, Lainey, Williams (2007-2012)

Remus et al. (2012)

Ogilvie & Lin (2004)

Greenberg (2009)

S.Ferraz-Mello (2012-2013)

Correa, Boué, Laskar & Rodríguez (2014)

- All of them depart in a larger or lesser extent from Darwin's theory.

Anelastic
Tide

New theory

Aim:

- Substitution of plugged
ad hoc lags by one **physical law**

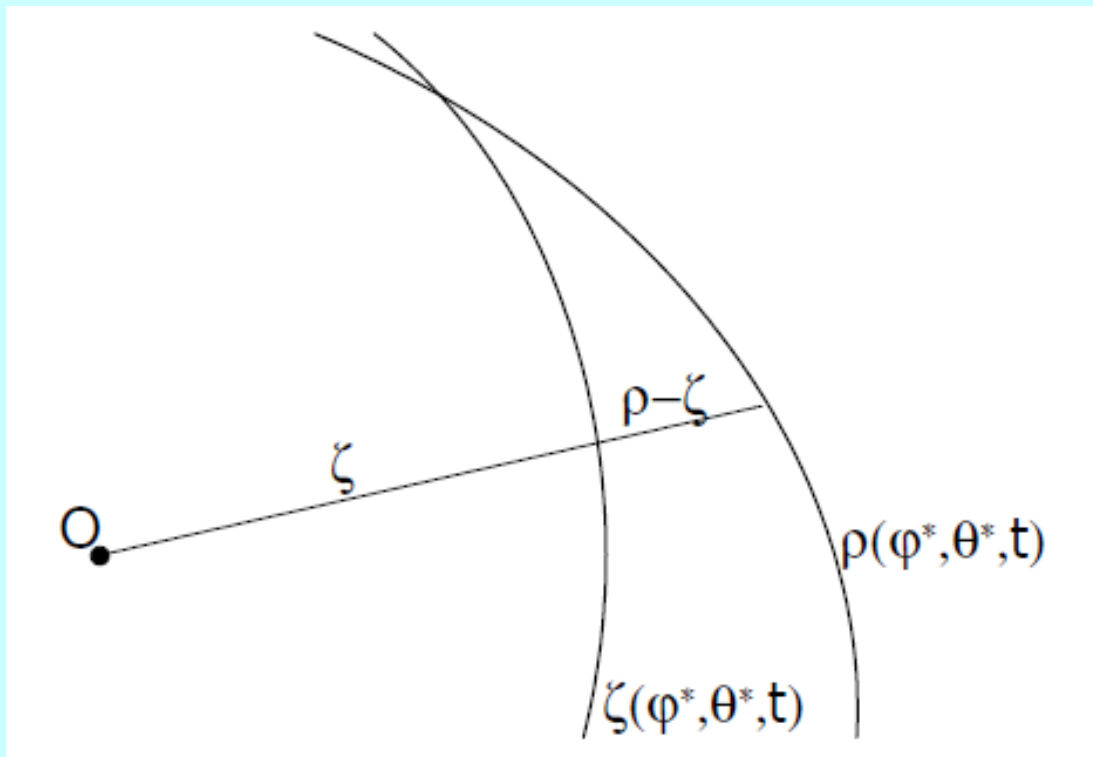
S.Ferraz-Mello

DDA 2012 (astro-ph 1205.3957)

Cel.Mech.Dyn.Astron.116,109(2013)

ANSATZ

$$\dot{\zeta} = \gamma(\rho - \zeta).$$



ζ Surface of the body

ρ Surface of instantaneous equilibrium (**VIRTUAL**)

$$\dot{\zeta} = \gamma(\rho - \zeta).$$



Newtonian
CREEP

$$\gamma = \frac{wR}{2\eta} = \frac{3gm}{8\pi R^2\eta}$$

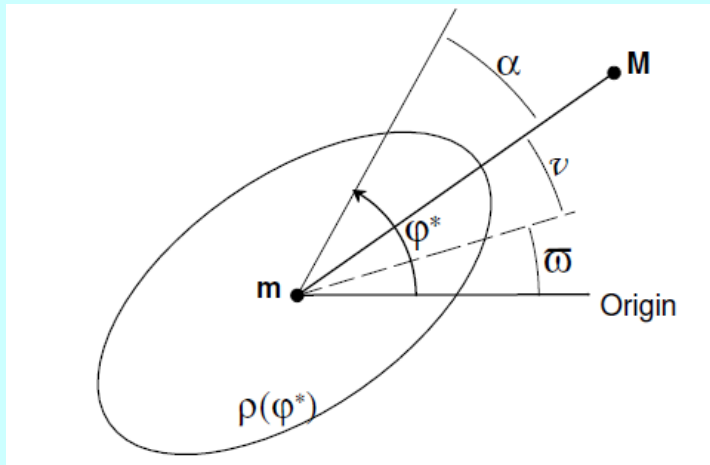
Relaxation factor
(critical frequency)

Approx. Solution of the
Navier-Stokes equation

Ref: Happel and Brenner, Low Reynolds number

Hydrodynamics, Kluwer, 1973 + **Darwin, 1879**

$$\dot{\zeta} + \gamma\zeta = \gamma R' + \frac{15\gamma R \sin^2 \theta^*}{8} \left(\frac{M}{m}\right) \left(\frac{R}{a}\right)^3 \left(\frac{a}{r}\right)^3 \cos(2\varphi^* - 2\varpi - 2v).$$



O.D.E. for $\zeta(t)$

$$\varpi = \varpi_0 - \Omega t.$$

Simplifications **(here!)**:

homogeneous bodies

equator = orbit plane; $\theta^* = \pi/2$

solution

$$\zeta = Ce^{-\gamma t} + R' + \frac{15\mu R \sin^2 \theta^*}{8} \left(\frac{M}{m}\right) \left(\frac{R}{a}\right)^3 \sum_{k=-N}^N \frac{E_k(e) \cos(2\bar{\alpha} + k\ell - \sigma_k)}{\sqrt{\mu^2 + (\nu + kn)^2}}$$

= Superposition of tidal bulges

Prolateness:

$$\epsilon_k = \frac{15}{4} E_k(e) \cos \sigma_k \left(\frac{M}{m}\right) \left(\frac{R}{a}\right)^3$$

$$E_k(e) = \frac{1}{2\pi\sqrt{1-e^2}} \int_0^{2\pi} \frac{a}{r} \cos(2v + (k-2)\ell) dv.$$

Cayley
functions

$$\sigma_k = \arctan \left(\frac{kn + \nu}{\gamma} \right).$$

Phase of the forced terms

$$\mathbf{F} = \text{grad } U \quad \rightarrow \quad dW/dt = \mathbf{F} \cdot \mathbf{v}$$

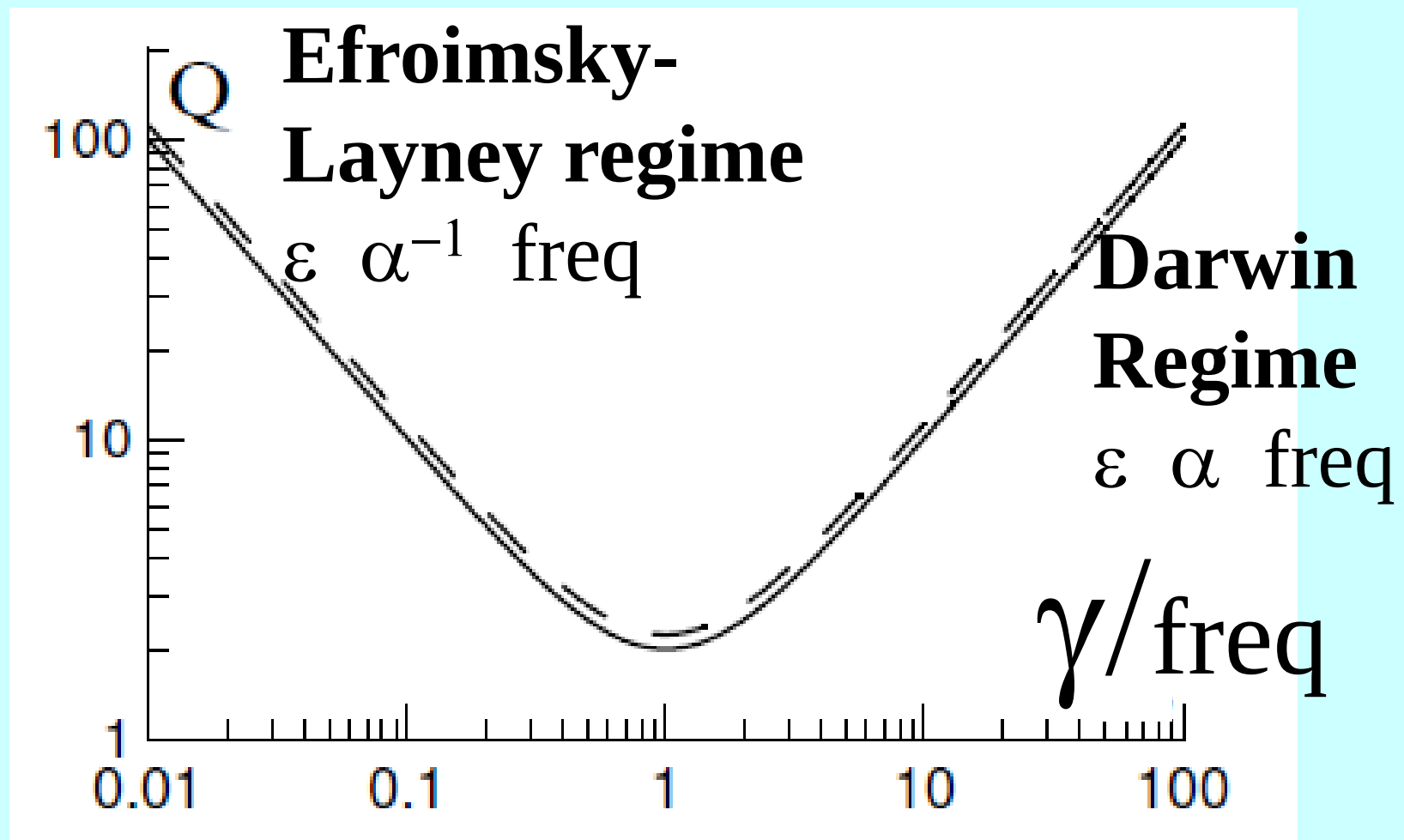
Energy dissipation

$$\langle \dot{W} \rangle_{\text{orb}} = - \frac{3k_f GM^2 R^5 n}{4a^6} \sum_{-N}^N (2 - k) E_k^2(e) \sin 2\sigma_k$$

Comparing to Darwinian linear theories:

$$\mathbf{Q} = \frac{k_2}{k_f} \left(\chi + \frac{1}{\chi} \right)$$

$$\chi = \text{Freq} / \gamma$$

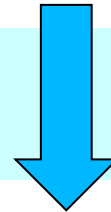


Free rotation $\text{freq} = \nu$ (semi-diurnal)

Pseudo-synchronous $\text{freq} = n$ (monthly)

Torque

$$M_2 = \sum_{k=-N}^N -\frac{2k_f GM m R^2 \epsilon_k}{5r^3} \sin(2v + (k-2)\ell - \sigma_k)$$



$$\dot{v} = -\frac{3GM\bar{\epsilon}_\rho}{2a^3} \sum_{k=2-N}^{2+N} E_{2,2-k} \sum_{j=-N-2+k}^{N-2+k} E_{2,2-k+j} (\sin 2\bar{\sigma}_k \cos j\ell + 2 \cos^2 \bar{\sigma}_k \sin j\ell)$$

Where $\boxed{v}$ is the semi-diurnal frequency = $2\Omega - 2\dot{\lambda}$

and

$$\sin 2\bar{\sigma}_k = \frac{2\gamma(\boxed{v} + (2-k)n)}{\gamma^2 + (\boxed{v} + (2-k)n)^2}$$

FIRST-order non-linear o.d.e.

SYNCHRONIZATION

Simulations
near $v=0$

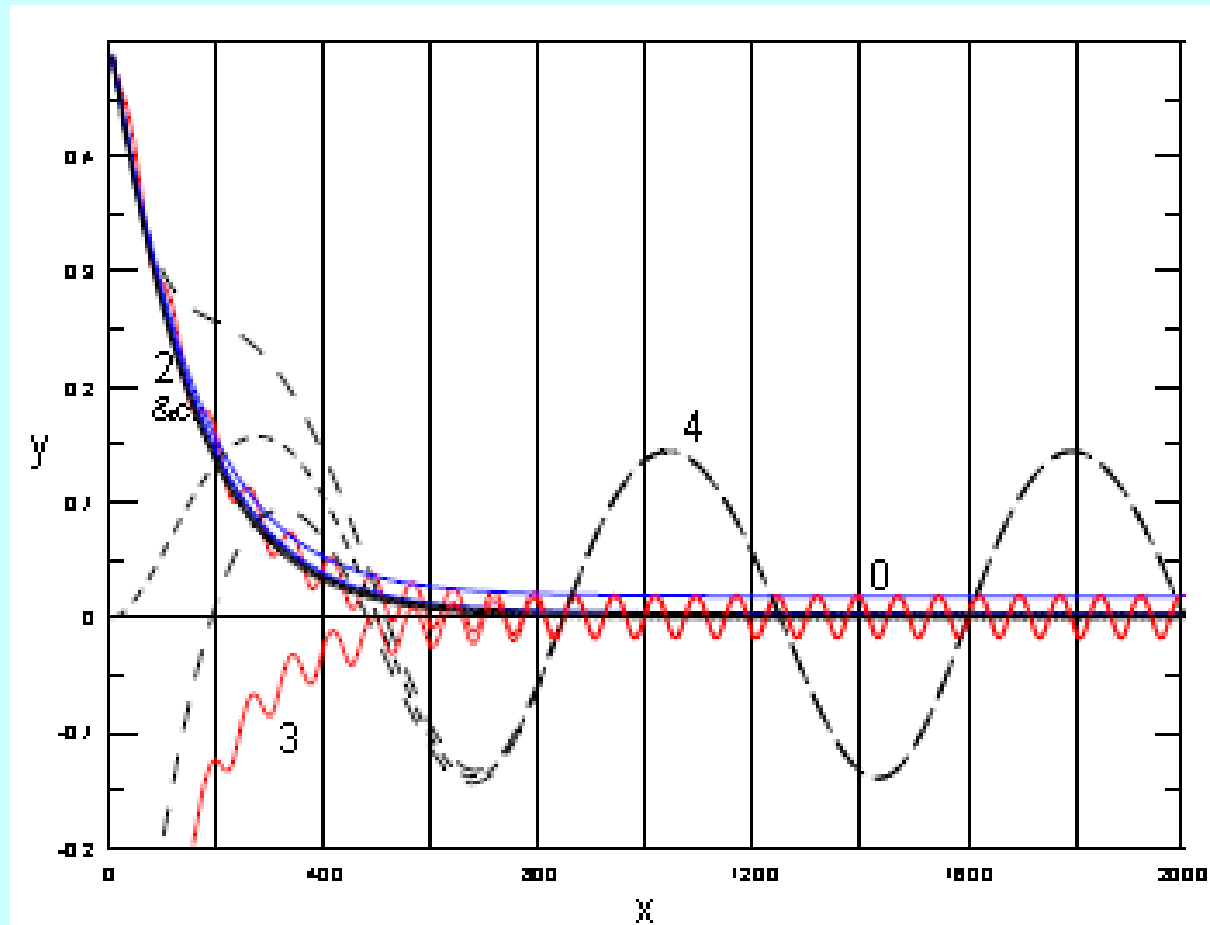
(normalized variables)

$$y = v/\gamma$$

$$x = (n/\gamma)(t - t_0)$$

Parameter

$$\log_{10} n/\gamma$$



$$\langle \dot{\Omega} \rangle = -\frac{45GM^2R^3}{16ma^6} \sum_{k=-N}^N E_k^2(e) \sin 2\sigma_k$$

$$\sin 2\sigma_k = \frac{2\gamma(\nu + kn)}{\gamma^2 + (\nu + kn)^2}.$$

• If $\langle \dot{\Omega} \rangle = 0$,

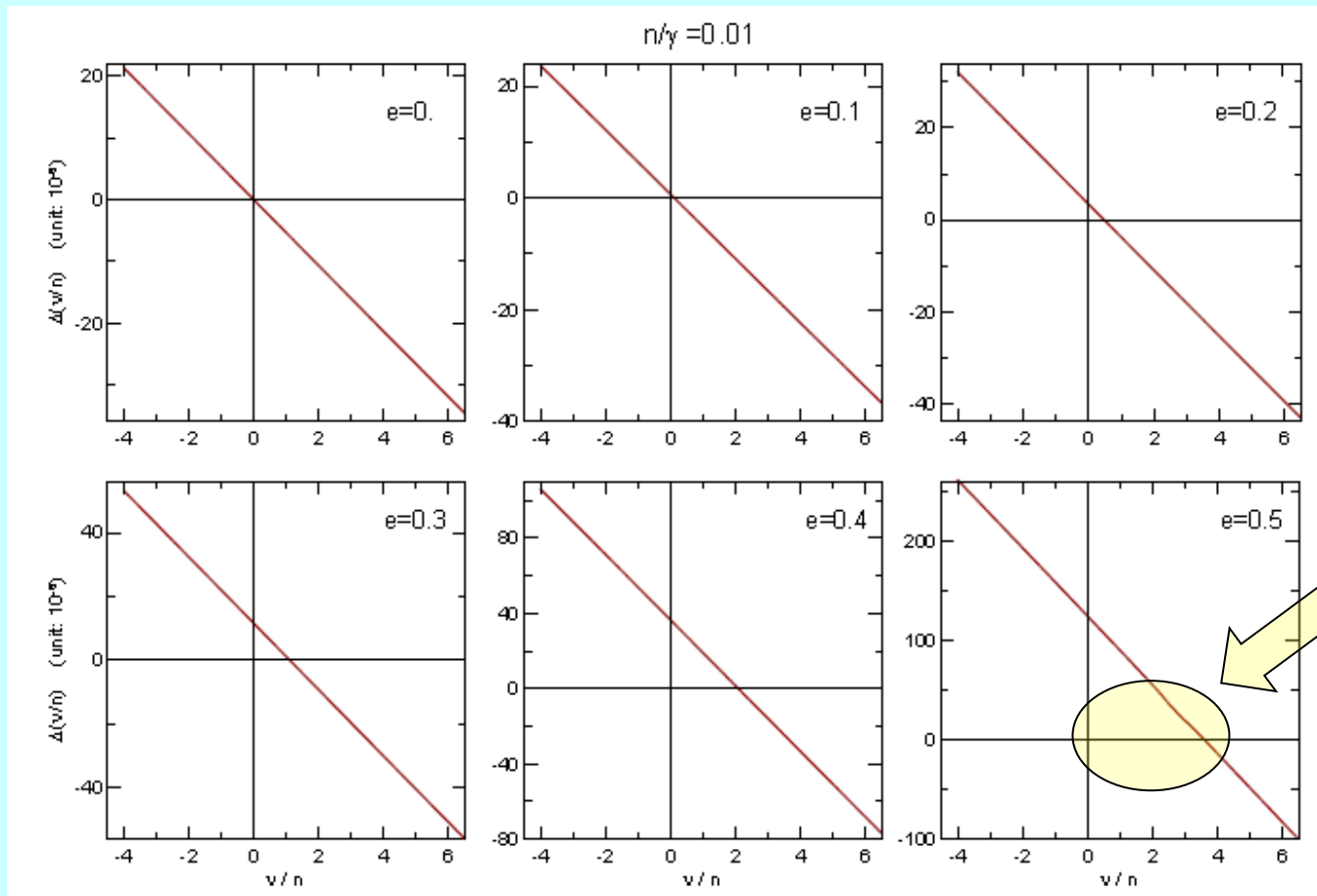
Stationary or pseudo-synchronous solution

$$\Omega = n + \frac{6n\gamma^2}{n^2 + \gamma^2} e^2 + 3n\gamma^2 \frac{226n^6 + 1453n^4\gamma^2 + 28n^2\gamma^4 + \gamma^6}{8(n^2 + \gamma^2)^3(4n^2 + \gamma^2)} e^4 + \mathcal{O}(e^6)$$

Limits: Efroimsky-Lainey $\gamma \ll n \rightarrow \Omega = n$
Darwin $\gamma \gg n \rightarrow \Omega = n(1+6e^2+..)$

Map

$$y(x) \rightarrow y(x+2\pi/\gamma) - y(x)$$

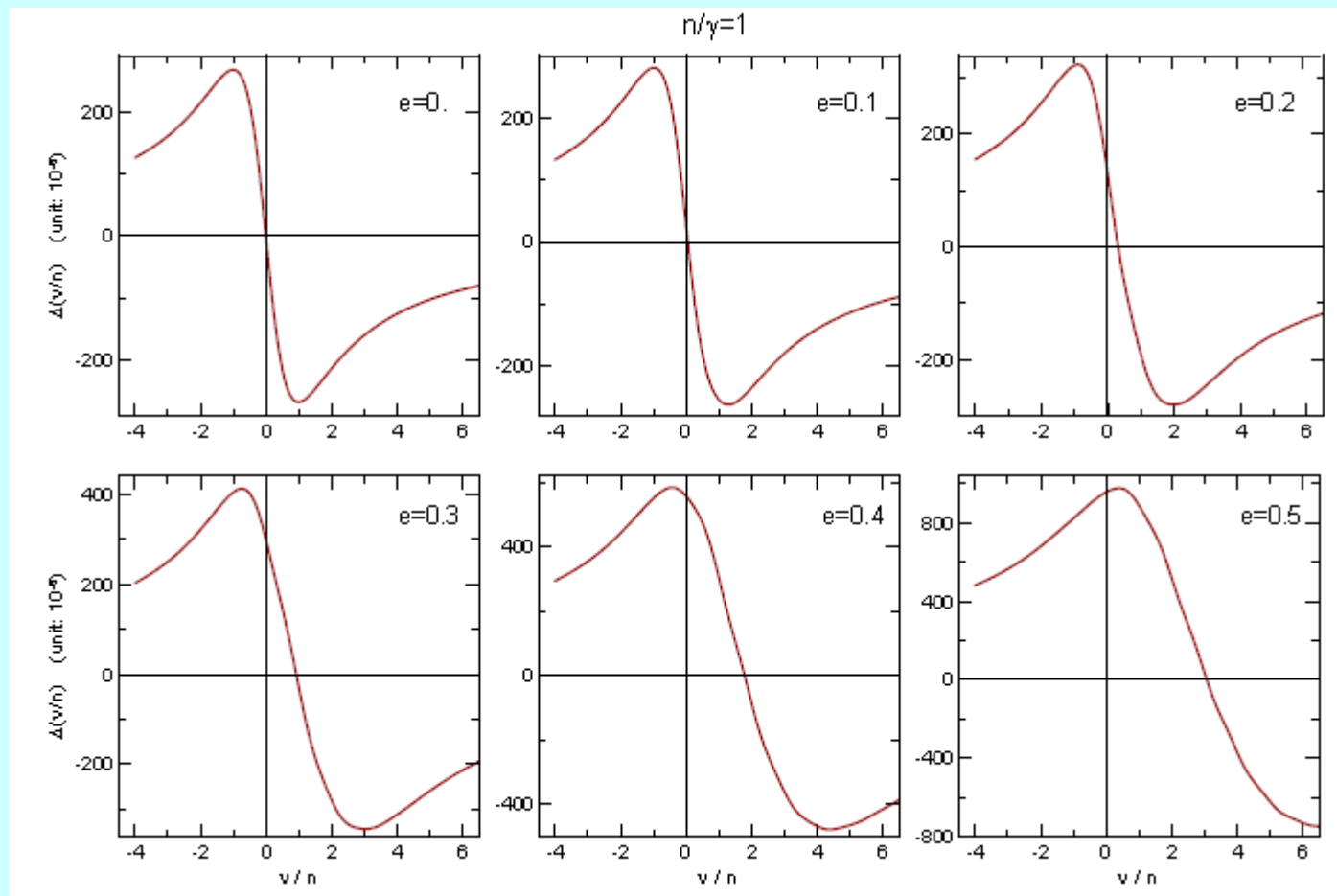


The intersections
are attractors

N.B.

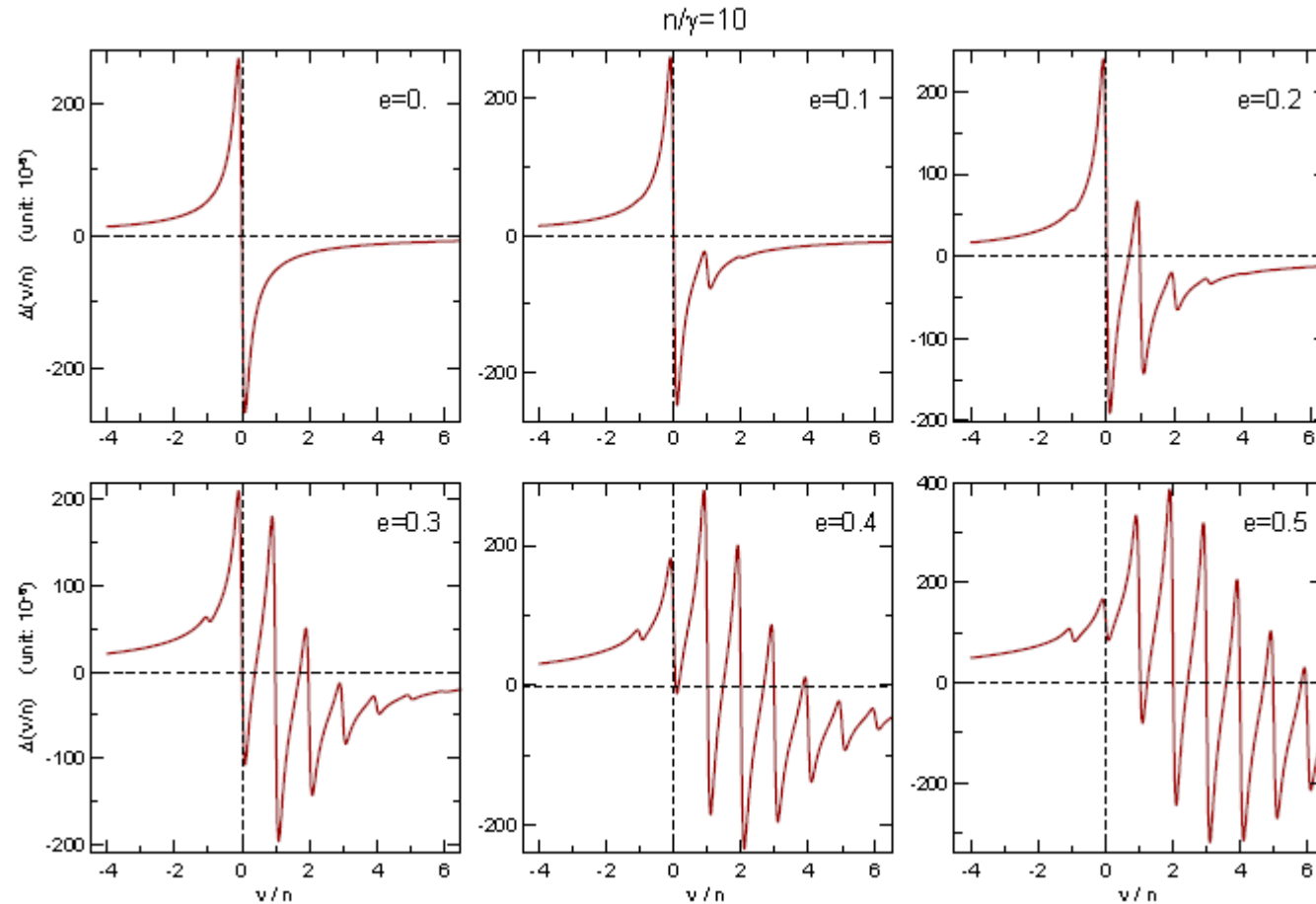
These attractors are
supersynchronous
 $\Omega = n + 6ne^2$

$\gamma \gg n$ (ex: stars, hot Jupiters)

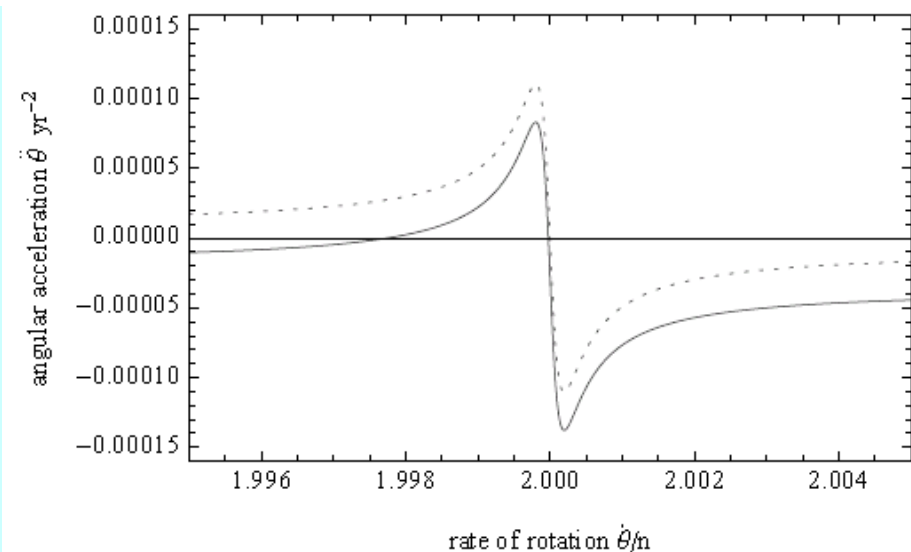


$$\gamma \sim n$$

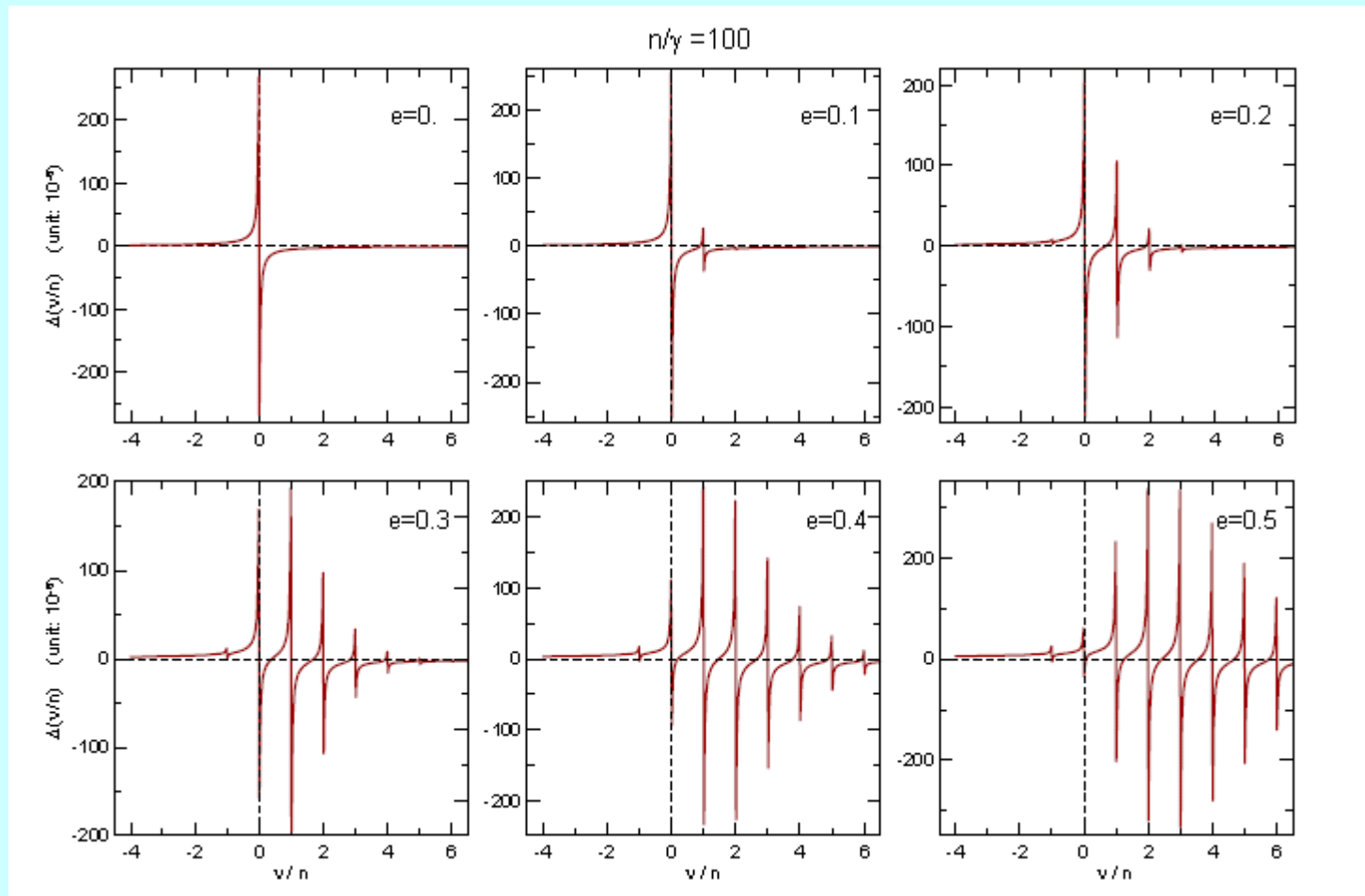
Ref: SFM, DDA 2014 and CeMDA (to be pub.)
Correia et al. A&A 2013



$\gamma = n/10$ (ex: Mercury,
distant satellites)
New attractors at $v=n, 2n, 3n, \dots$

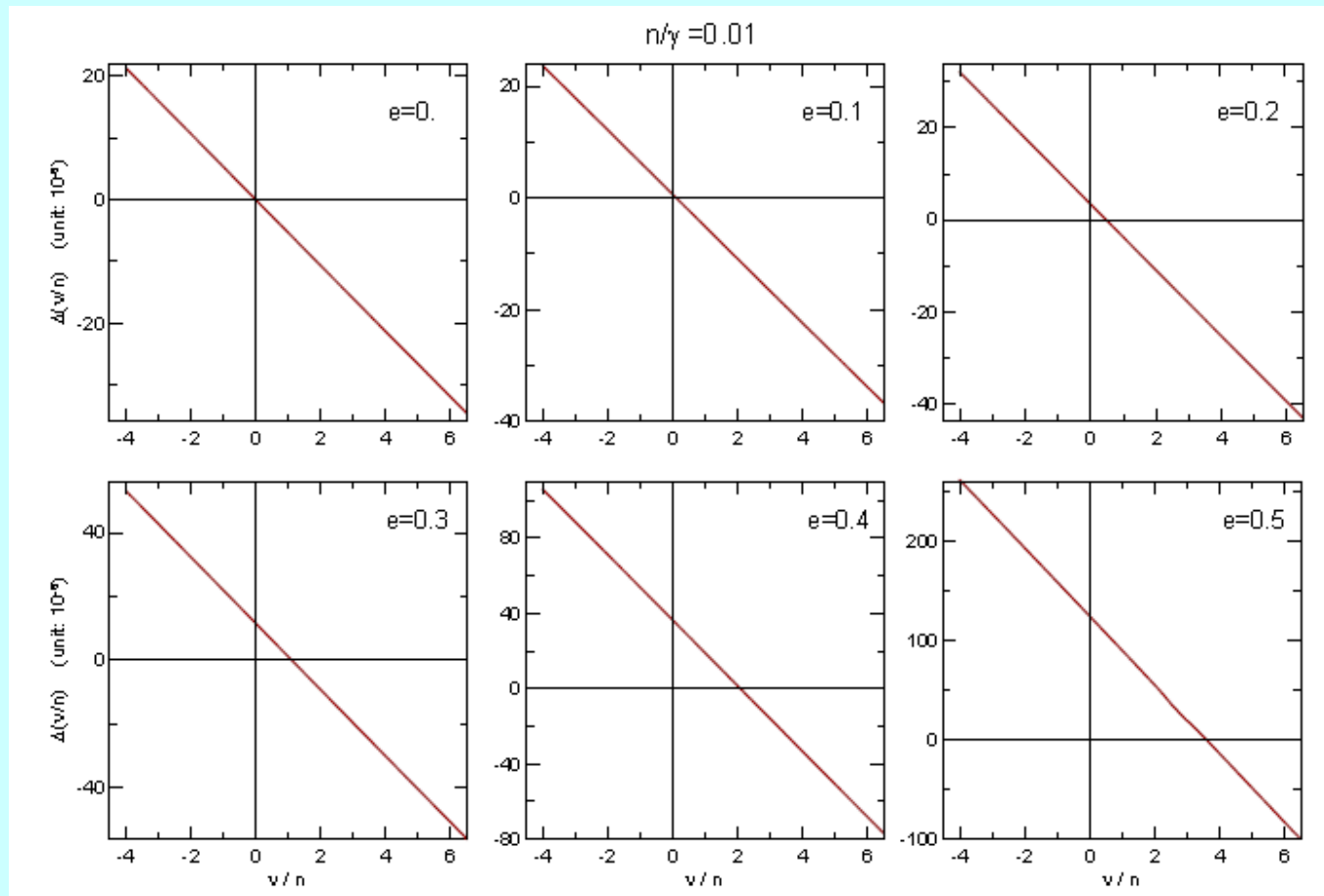


Makarov & Efroimsky, ApJ. 2013



$\gamma \ll n$ (ex: Moon, Titan)
 attractors at $v = -n, 0, n, 2n, 3n, \dots$
 $\Omega = n/2, n, 3n/2, 2n, 5n/2, \text{ etc.}$

STARS (high γ)



Then, for solar-type stars

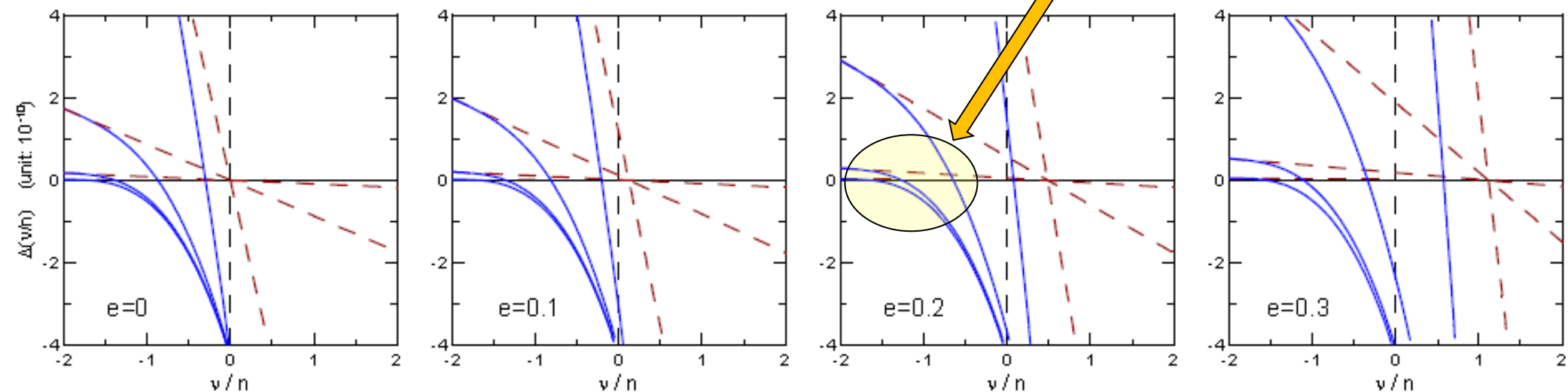
$$\dot{\Omega} = -f_P B_W \Omega^3$$

where

$$B_W = 2.7 \times 10^{47} \frac{1}{C} \sqrt{\left(\frac{R}{R_\odot} \frac{M_\odot}{m} \right)} \quad (\text{cgs units})$$

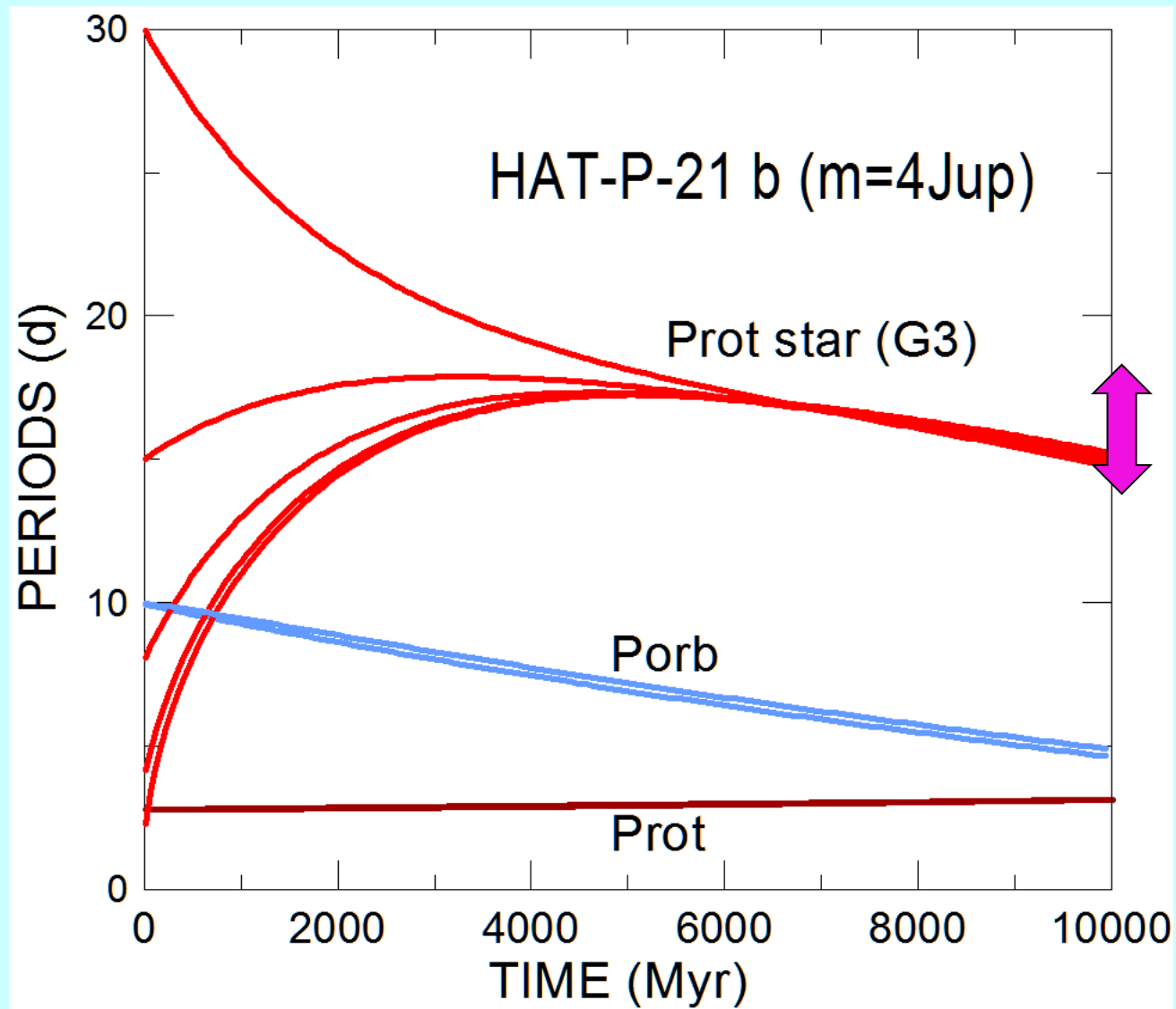
and $0 < f_P < 1$. **THEN**

subsynchronous
attractors



curves: $n/\gamma = 10^{-6}$ to 10^{-3}

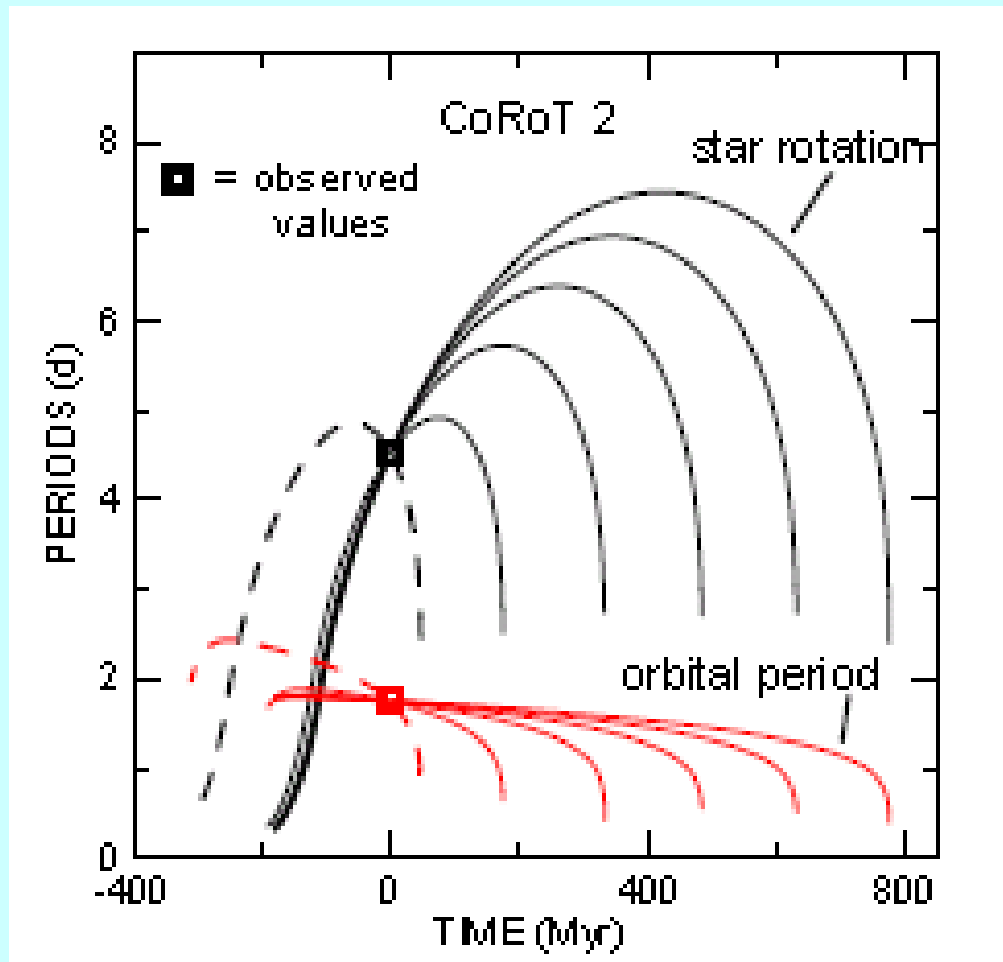
[brown: $f=0$; blue: $f=1$]



$$\gamma=25.5\text{s}^{-1}$$
$$e=0.228$$

Observed
value

CoRoT 2: A young star

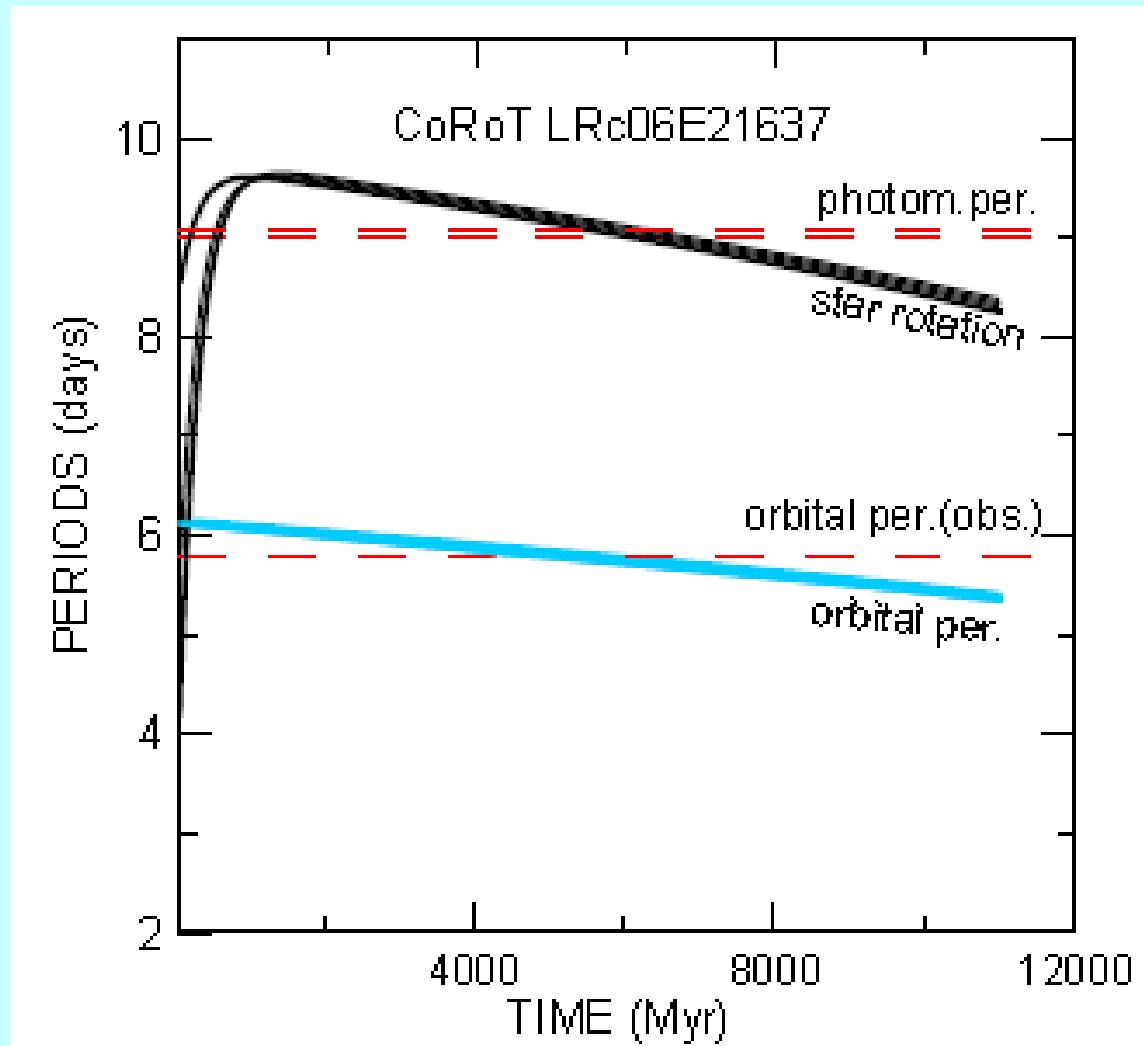


$$m_{pl}=3.3 \text{ jup}$$
$$\gamma=20-100 \text{ s}^{-1}$$

Result:
Age < 200 Myr

SFM et al. Astrophys. J. (in press)

CoRoT 33: A paradigm



$$m_{\text{comp}} = 62 \text{ jup}$$
$$\gamma = 55 \text{ s}^{-1}$$

SFM et al. Astrophys. J. (in press)

Two interesting F star cases:

CoRoT 15b BD (m=63.3 Jup) around a F7V star

Orbital period: 3.06 d

Star rotation: 2.9 – 3.1 d

KELT 1b BD (=27.4 Jup) around a F5 star

Orbital period: 1.217 d

Star rotation: $1.348 \pm 0.4 \sin I$

THE END

THANKS

MERCI BEAUCOUP

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